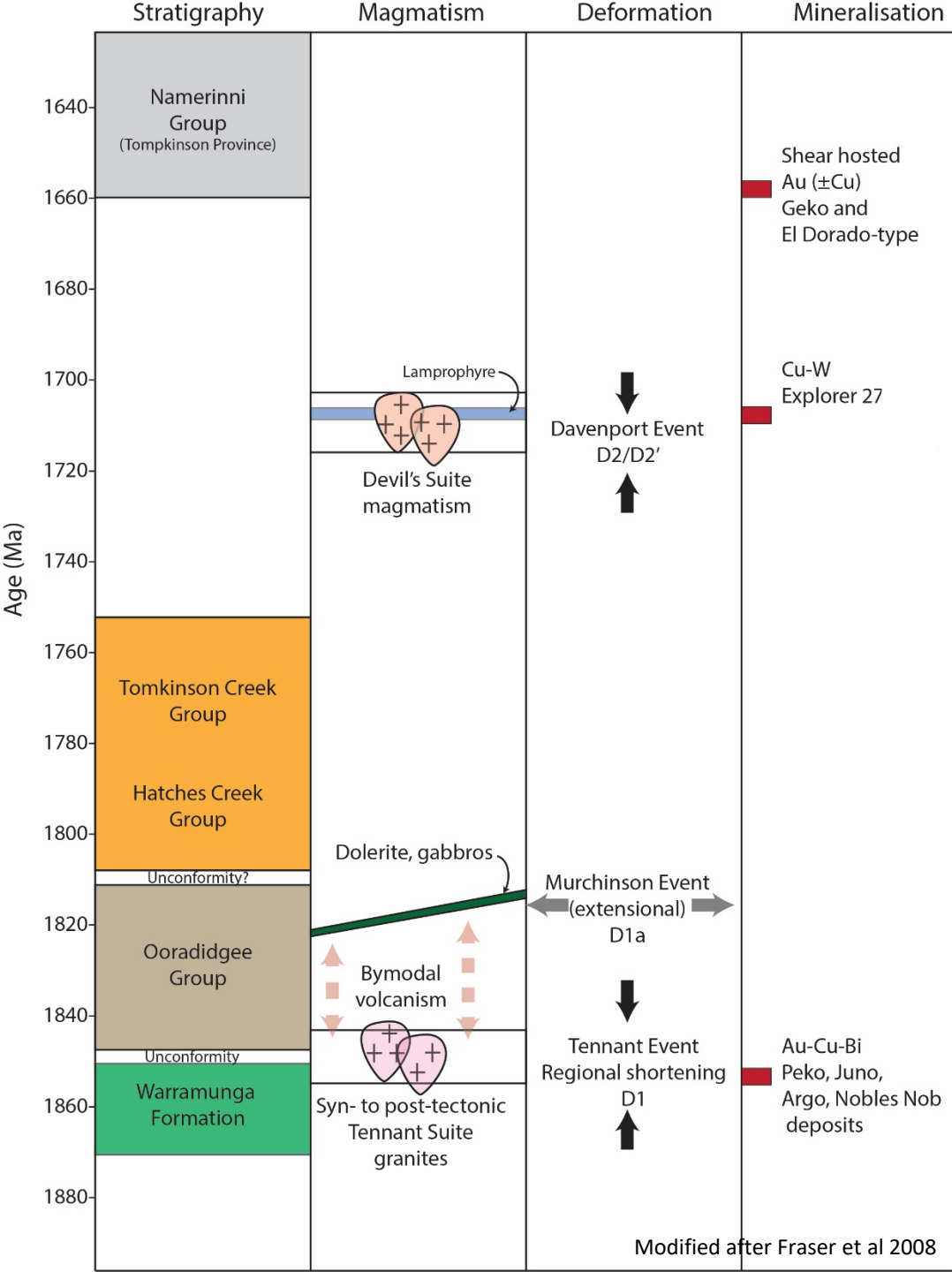


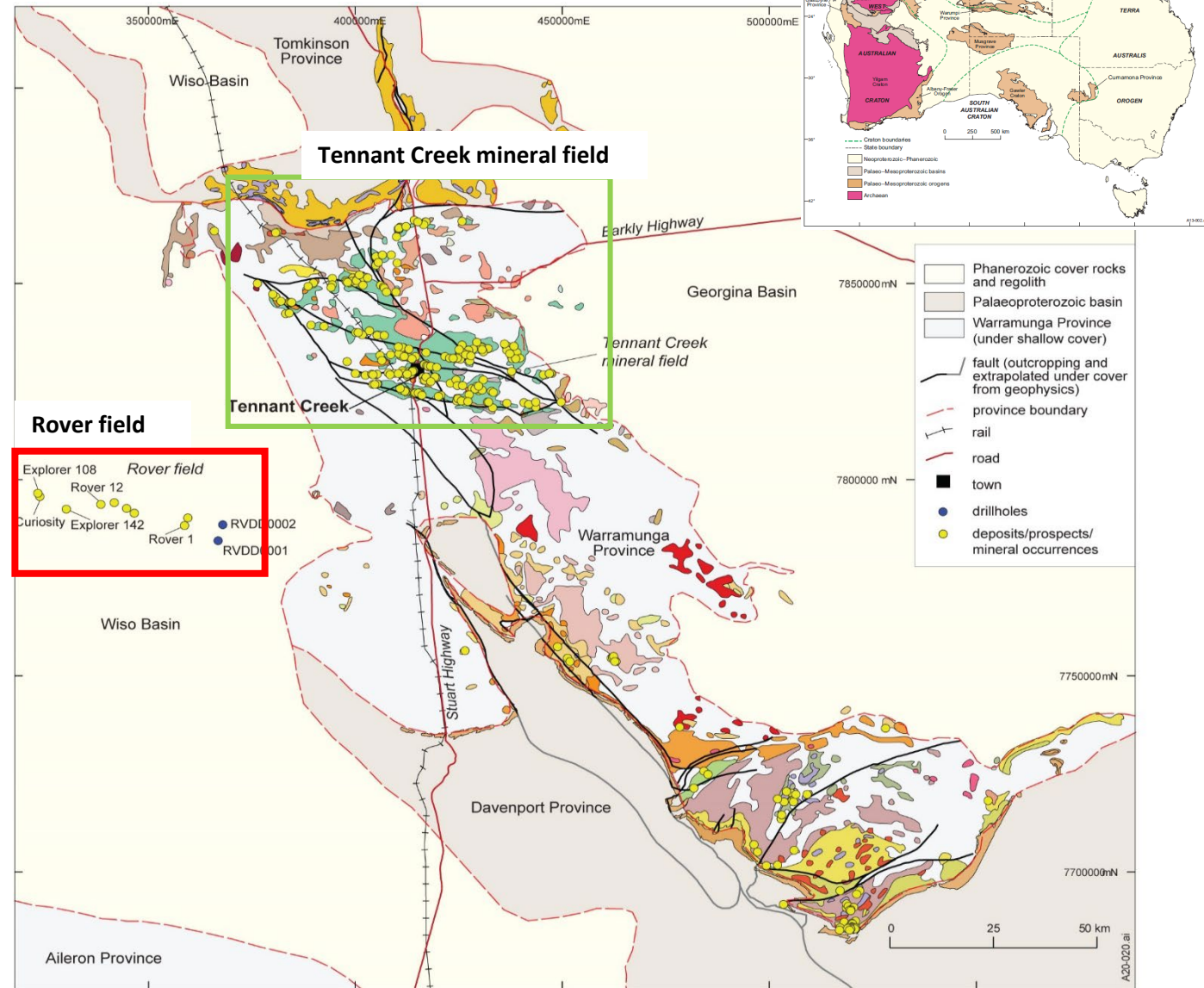
# The Rover field: Insights on stratigraphy, age and base metal mineralisation

Pablo Farias<sup>1</sup>, Andrew Cross<sup>2</sup>, David Huston<sup>2</sup>





## Location and geological background



Questions:

What is the lithology and age of the covered basement in the Rover field?

What kind of deposits are hosted in the Rover field? Are they the same than in the Tennant Creek mineral field?

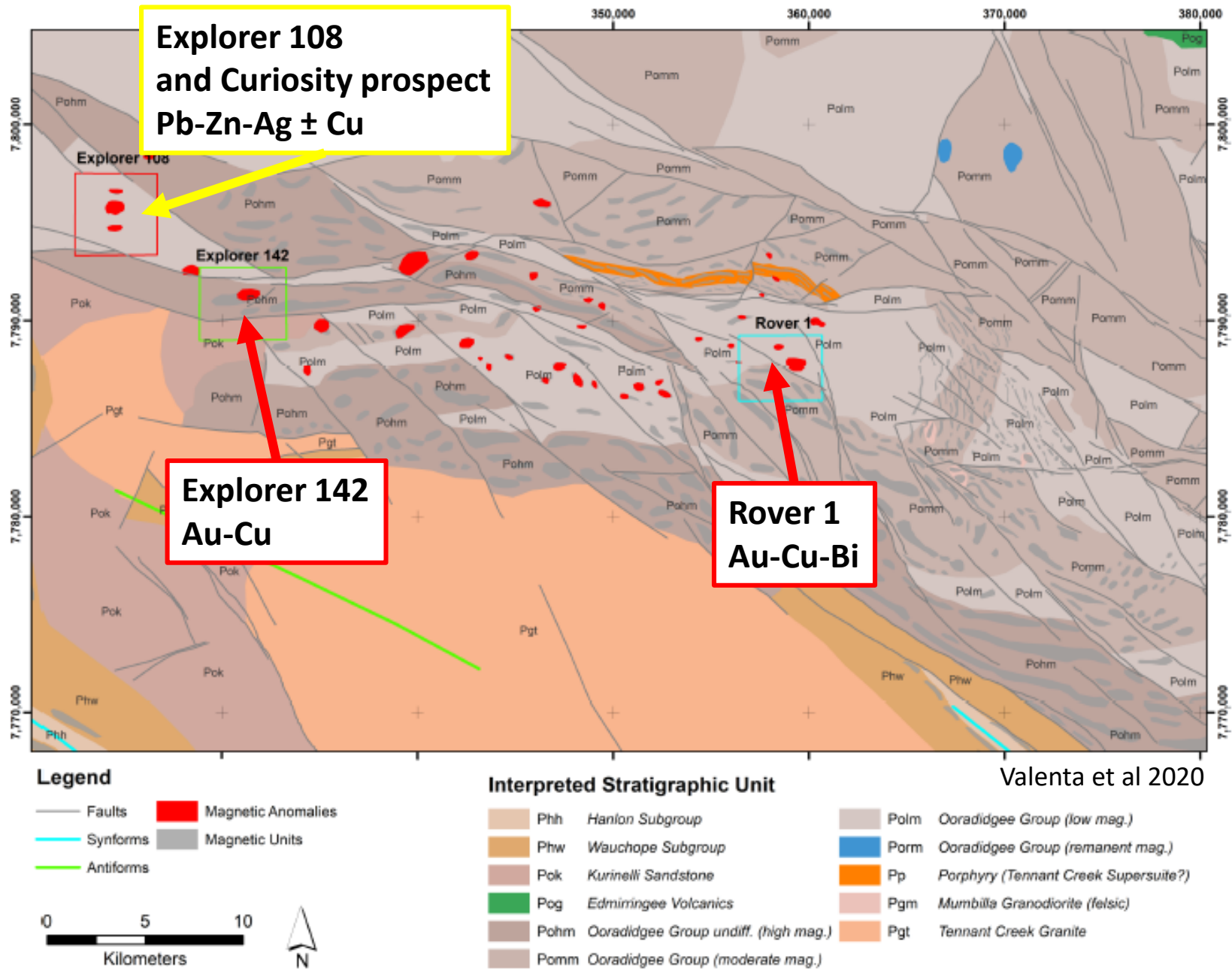
What is the age of these deposits?

# Rover field

## Mineral deposits and prospect

**Explorer 108**  
JORC 2012 mineral resource of 11.9Mt at 11.1g/t Ag, 2.0% Pb, 3.2% Zn for 4.25Moz silver, 237kt lead, 385kt zinc, above a lower cut-off grade of 2.5% combined Pb-Zn

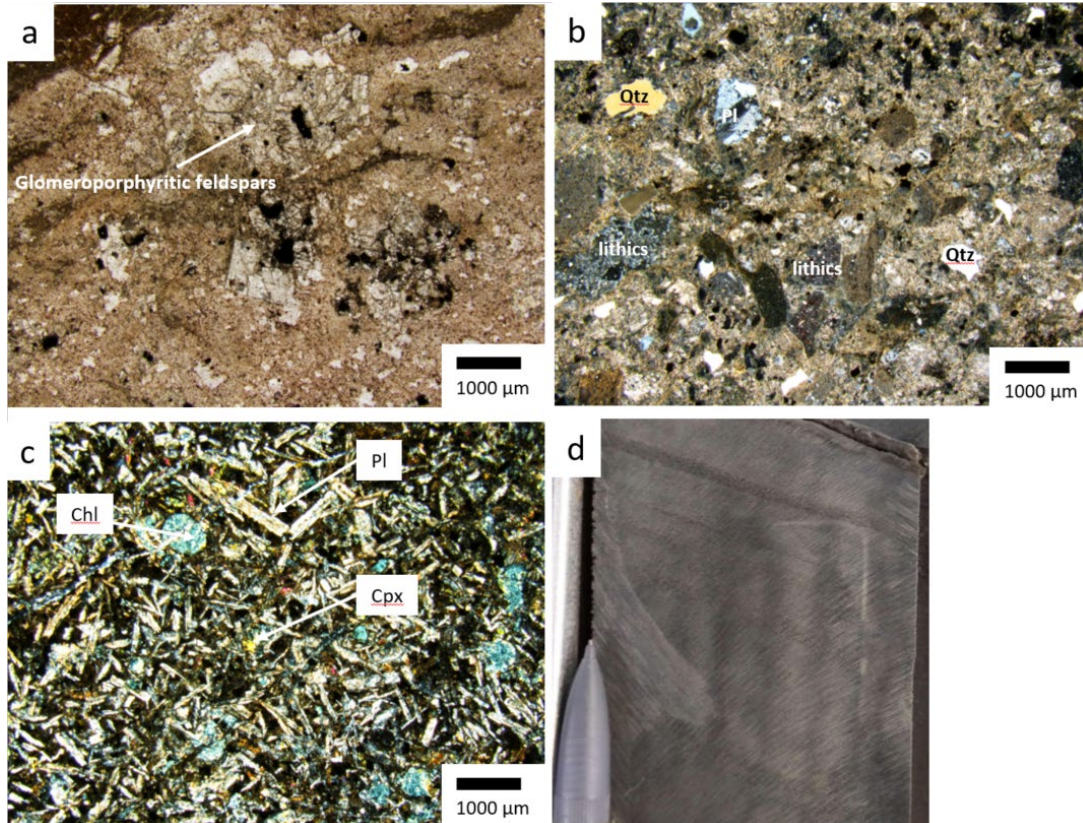
**Rover 1**  
JORC 2012 mineral resource of 6.9 Mt at 1.74 g/t Au, 2.07 g/t Ag, 1.2 % Cu, 0.14 % Bi, 0.06 % Co for 386 koz gold, 459koz silver, 83kt copper, 9.4kt bismuth and 4.1kt cobalt



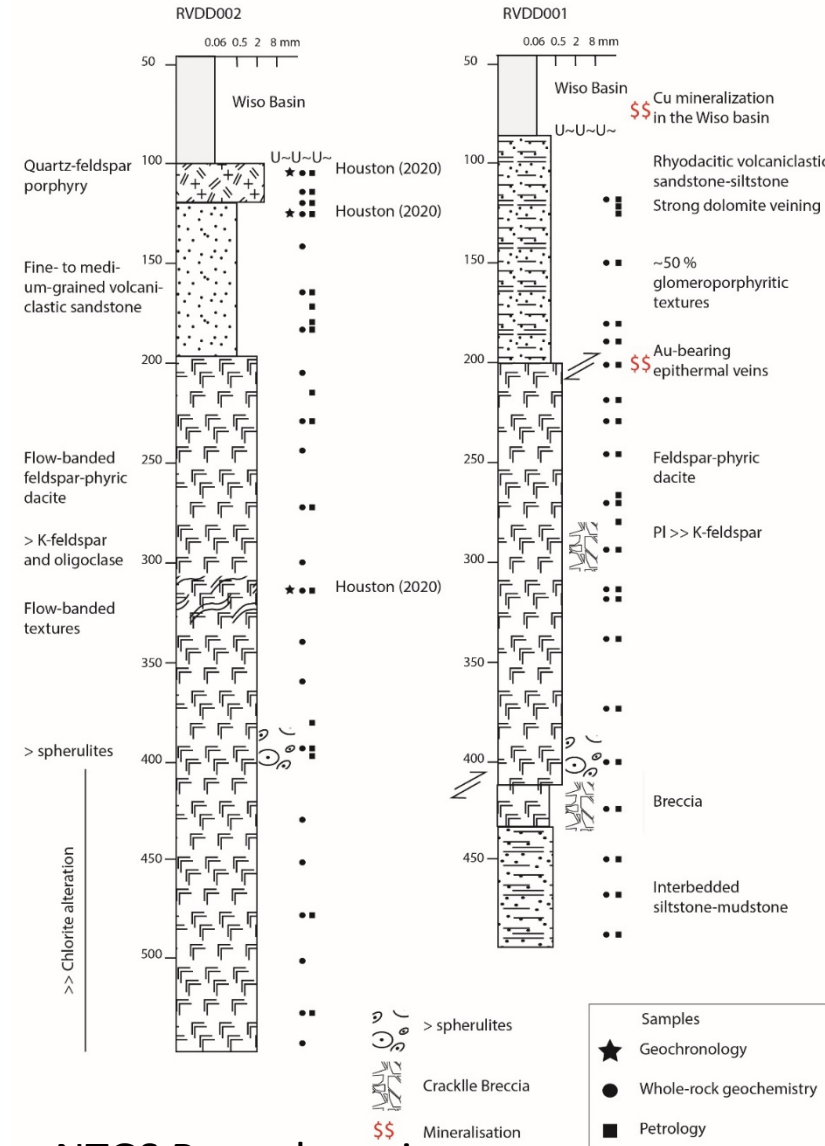




# Petrography of Rover field



- 20 drillholes logged and sampled
- Over 200 samples for petrology and whole-rock geochemistry



NTGS Record coming soon

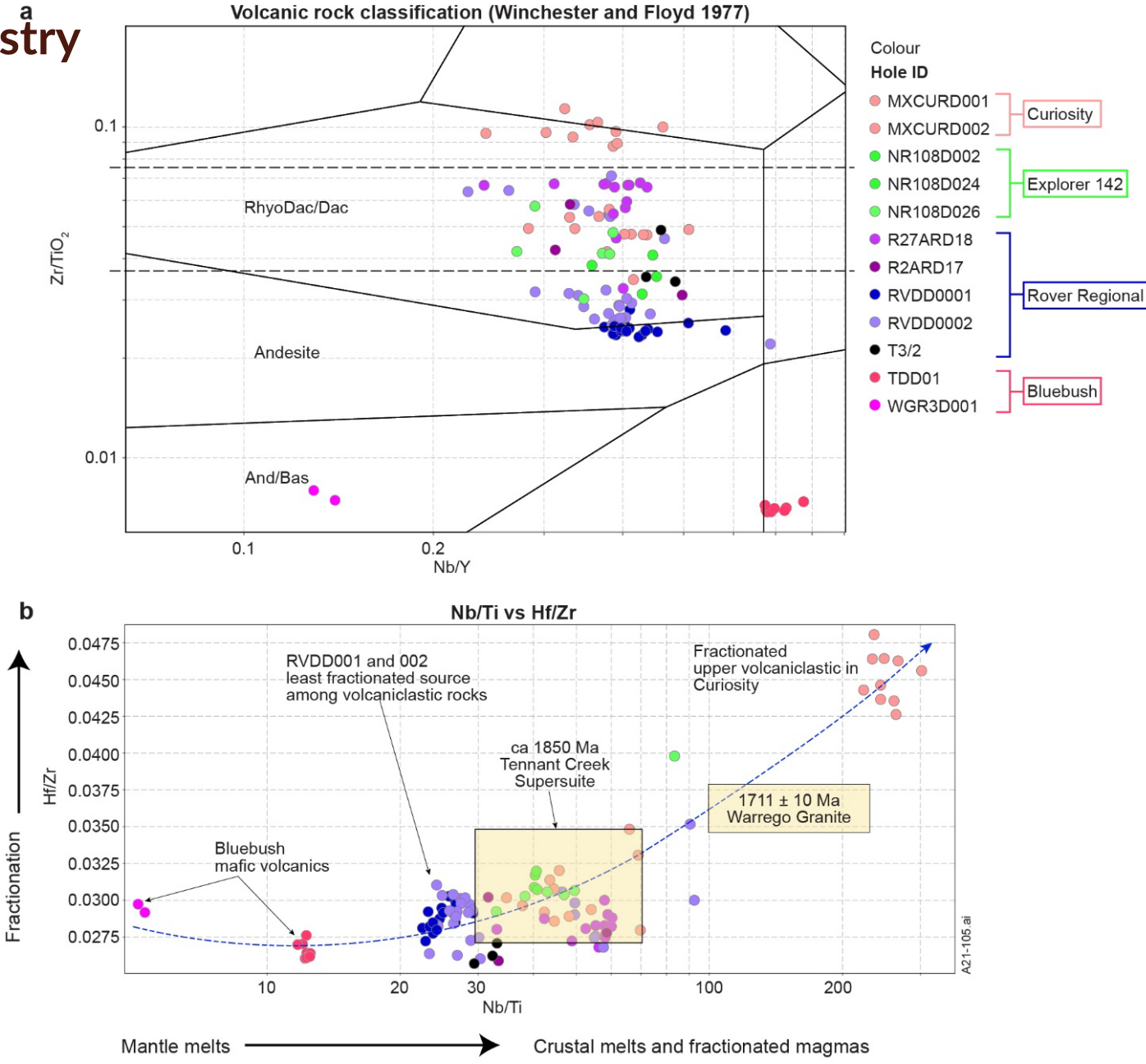
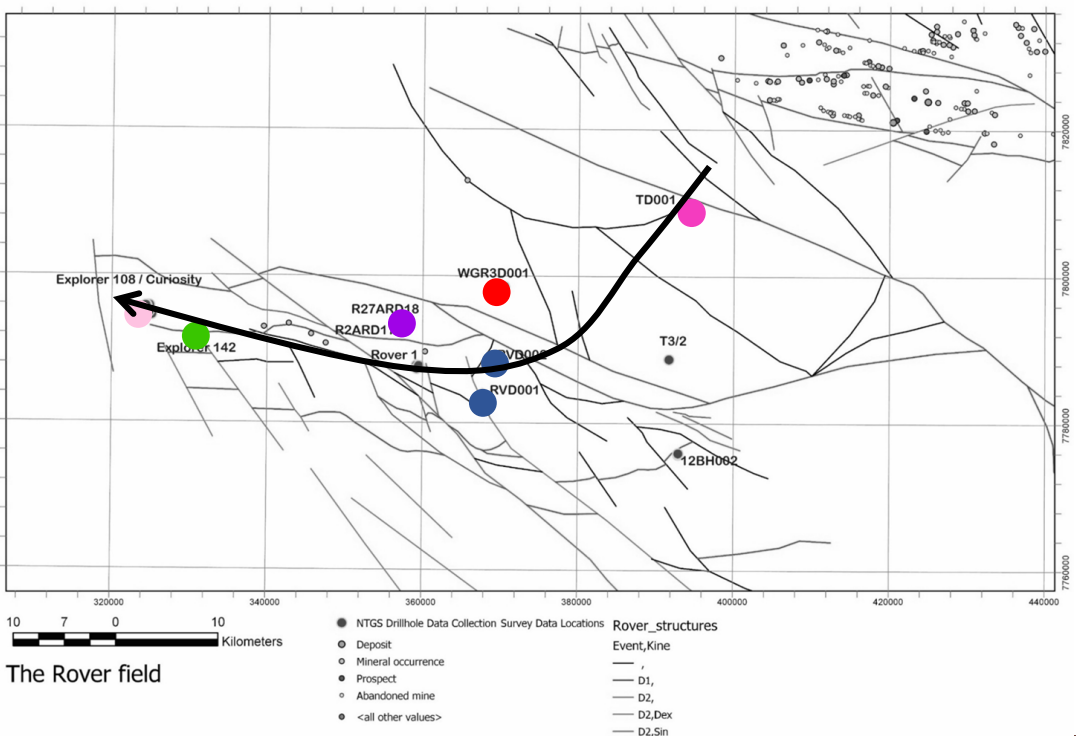
Four major lithological groups:

- Coherent dacitic and andesitic volcanic rocks
- Immature rhyolitic volcanoclastic rocks
- Turbiditic siliciclastic rocks
- Mafic intrusive and extrusive rocks
- Minor quartzofeldspathic porphyritic intrusions



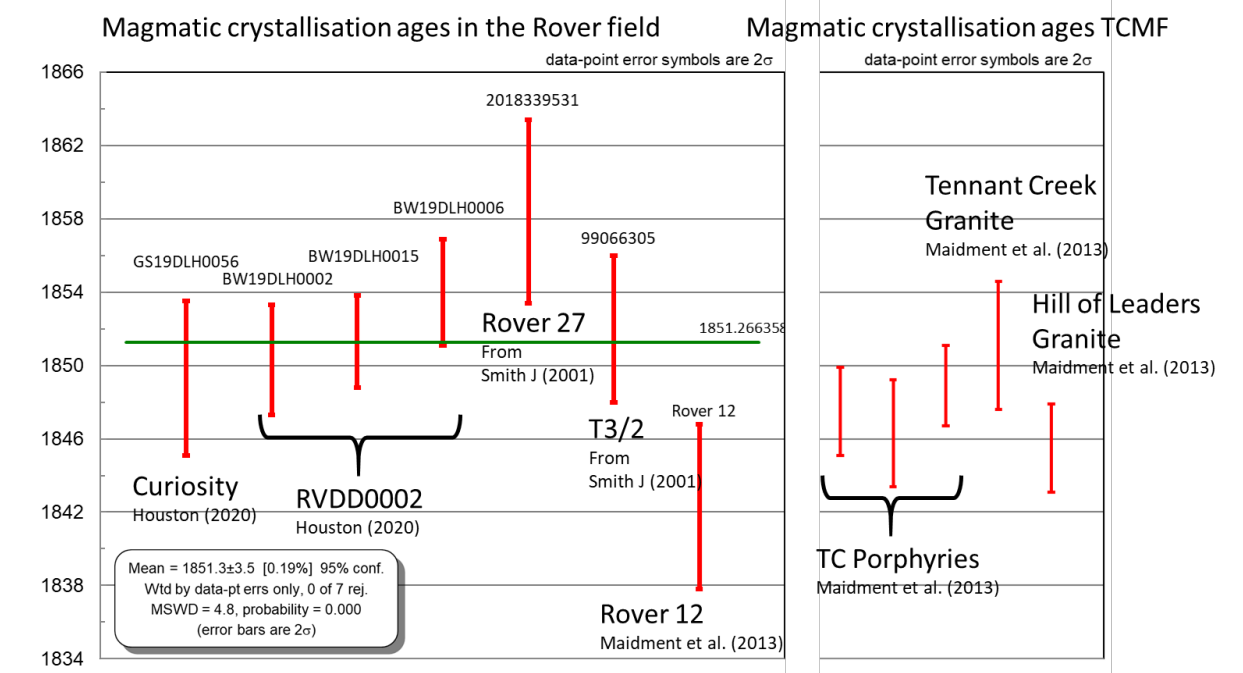
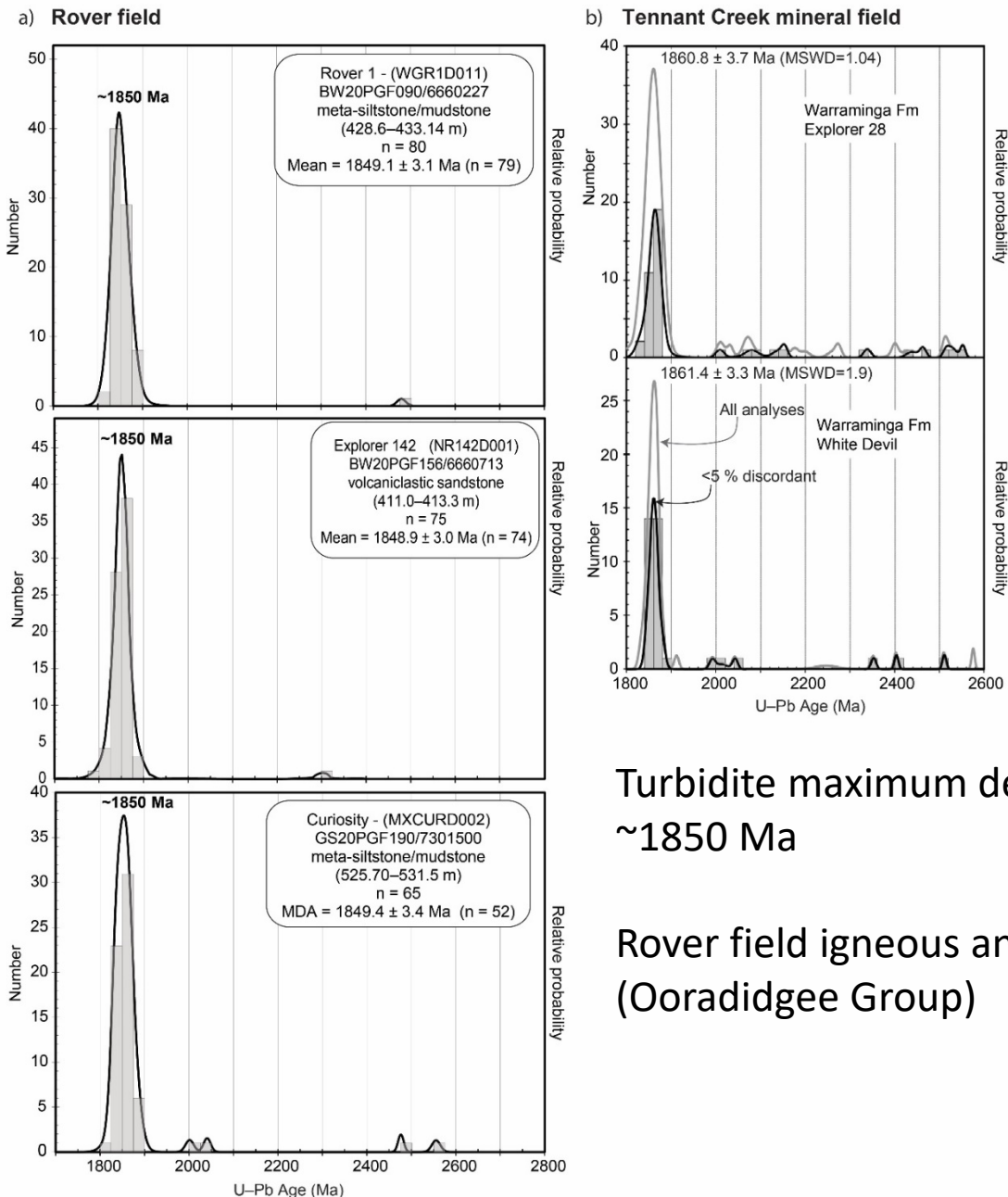
# Volcanic and volcanoclastic rocks geochemistry

- Mostly of intermediate compositions (Rhyodacitic)
- **Mafic** and primitive compositions in the north
- **Andesitic** compositions in the east
- **Rhyolitic and fractionated** compositions in the west



# Rover field host rocks geochronology

## Zircon U–Pb (MDA and crystallisation ages)

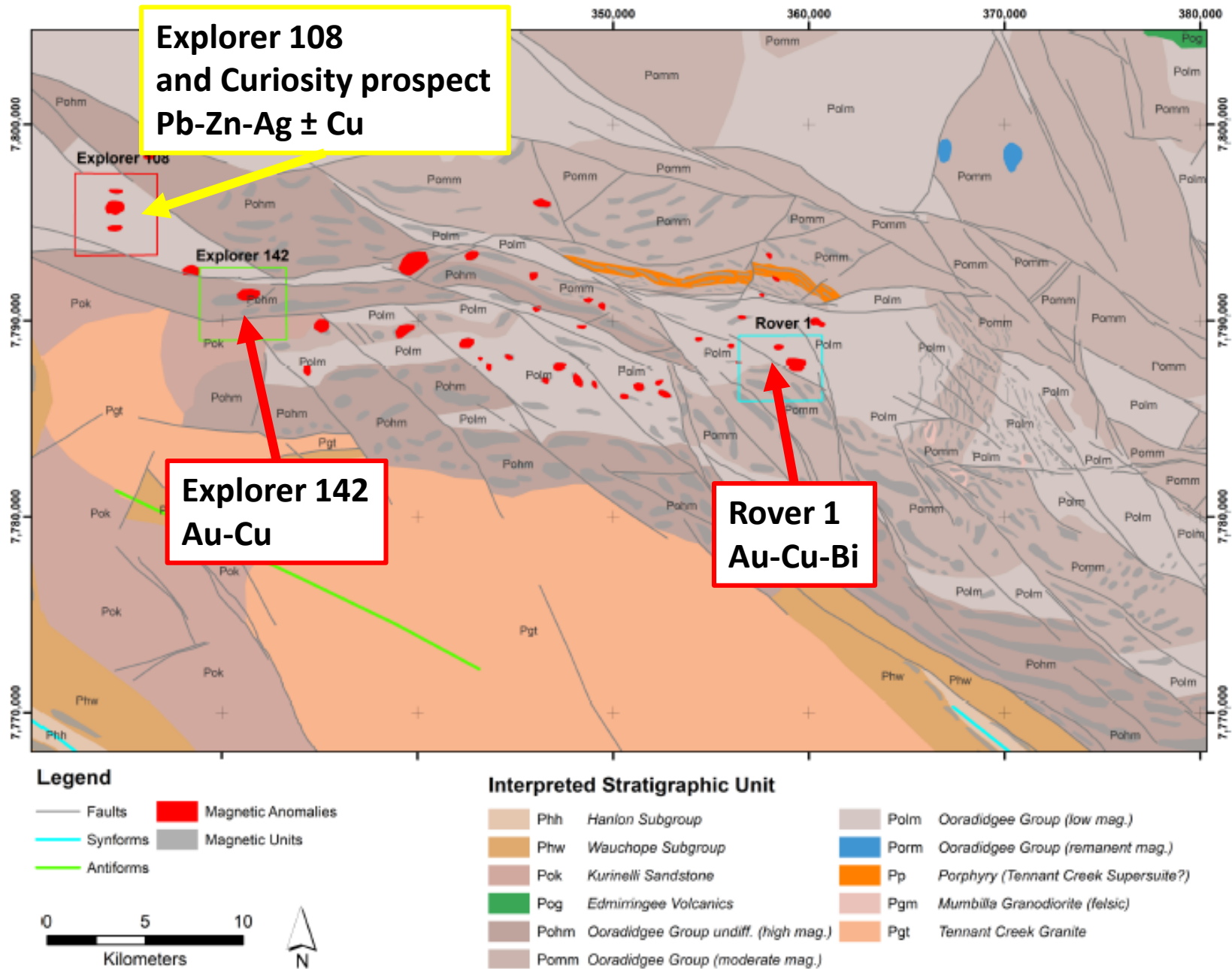


Turbidite maximum depositional ages a(MDA) and volcanic rocks crystallisation ages:  
~1850 Ma

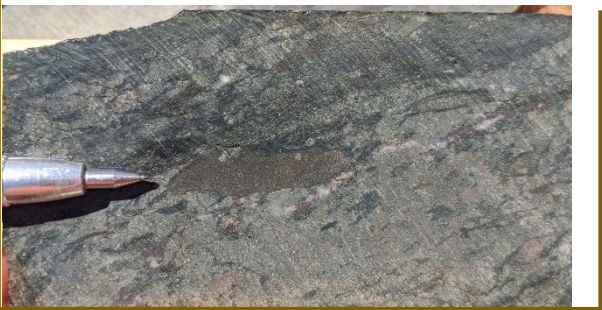
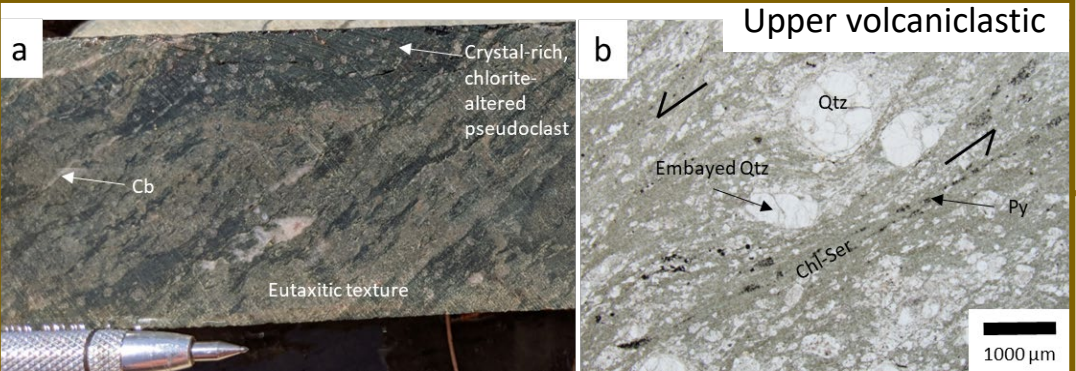
Rover field igneous and sedimentary rocks are equivalents to the **Yungkulungu Fm**  
(Ooradidgee Group)



# Explorer 108 - Curiosity

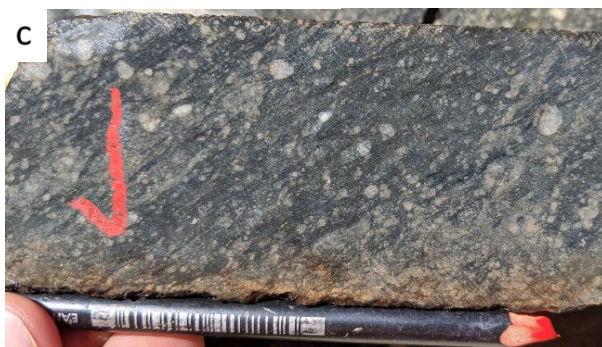


# Explorer 108 – Curiosity host rocks

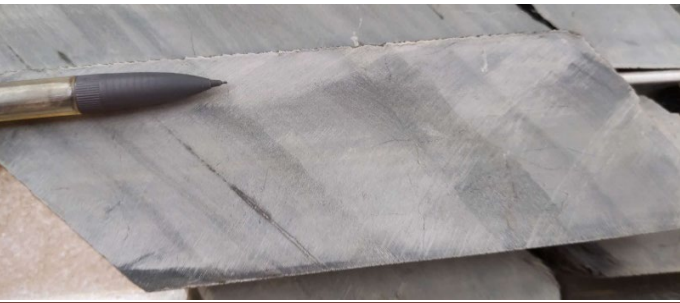
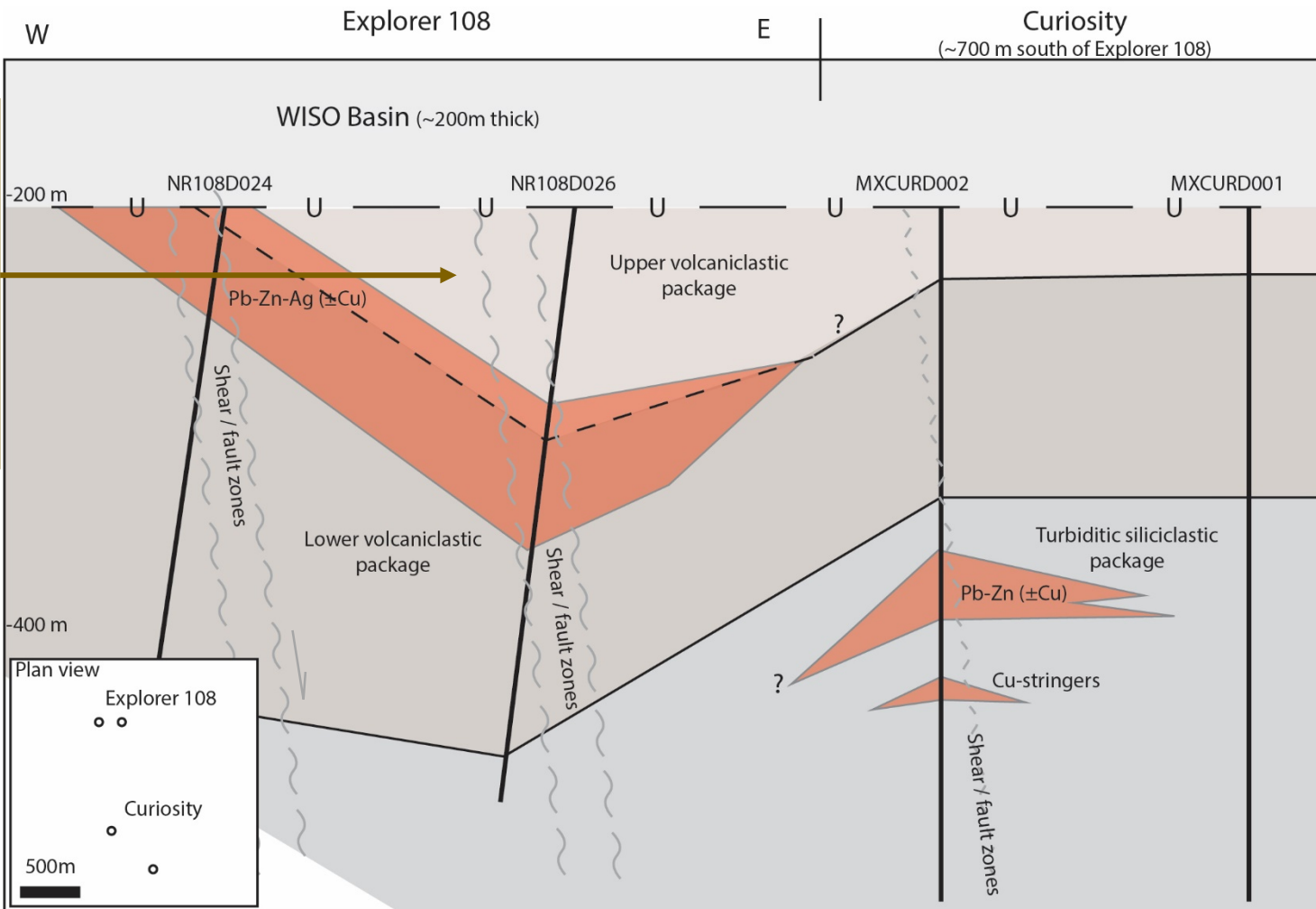


Foliated, sheared volcaniclastic breccia with clasts of the lower volcaniclastic package, and fine-grained volcaniclastic sediment matrix

## Lower volcaniclastic



Foliated, chlorite-altered massive, matrix supported, crystal-rich volcaniclastic sandstone



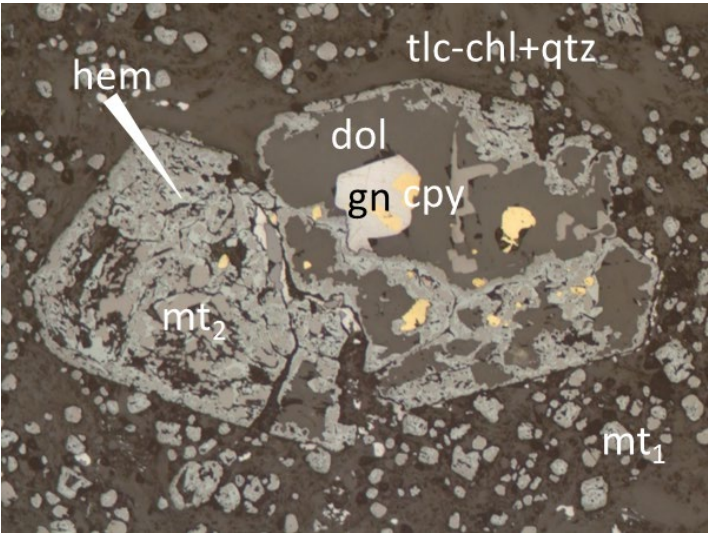
Fine-grained, laminated metasandstone-siltstone (Turbidite)



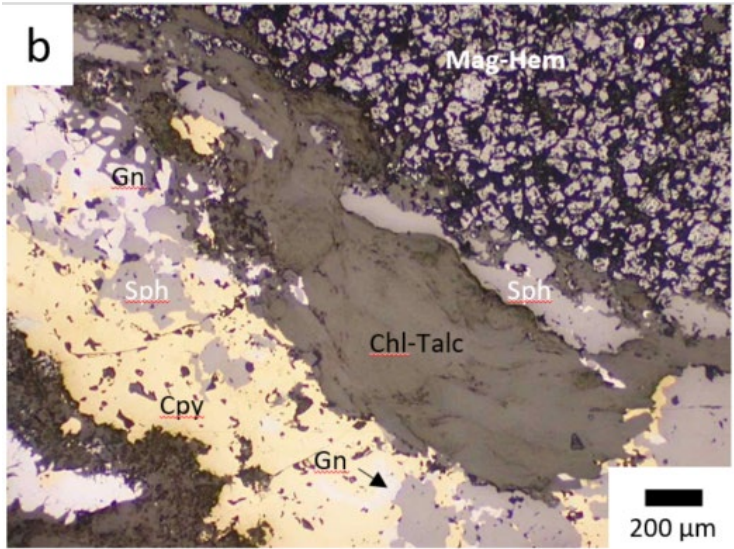
# Curiosity mineralisation



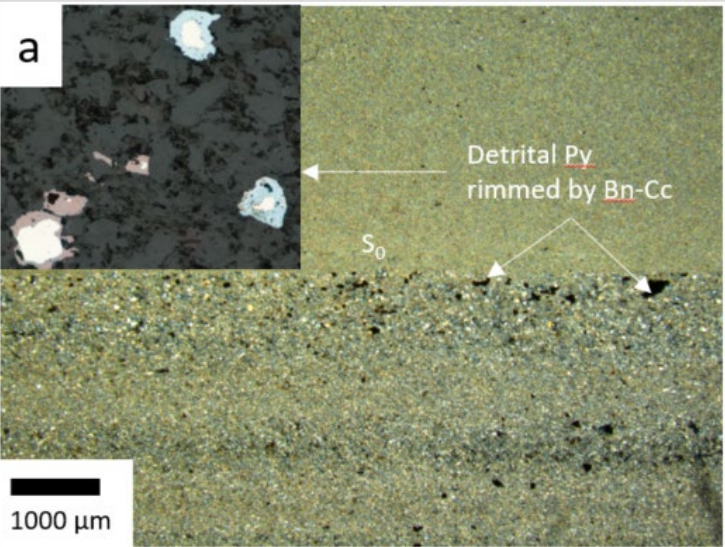
Folded massive to semi-massive sulfides (py-cpy-sph-gn + cb-mt-chl-tlc gangue)



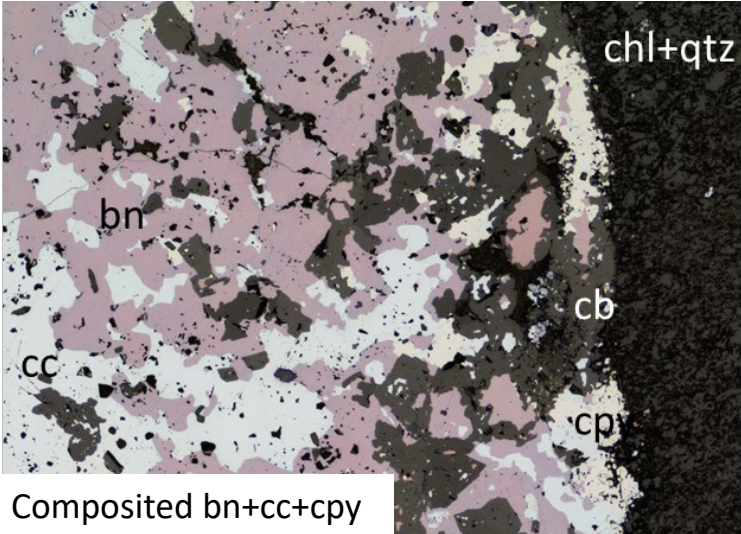
Mineralised zone: magnetite 1 and 2  
Magnetite replaced by Hematite



Cpy-Sph-Gn in Chl-rich shear planes



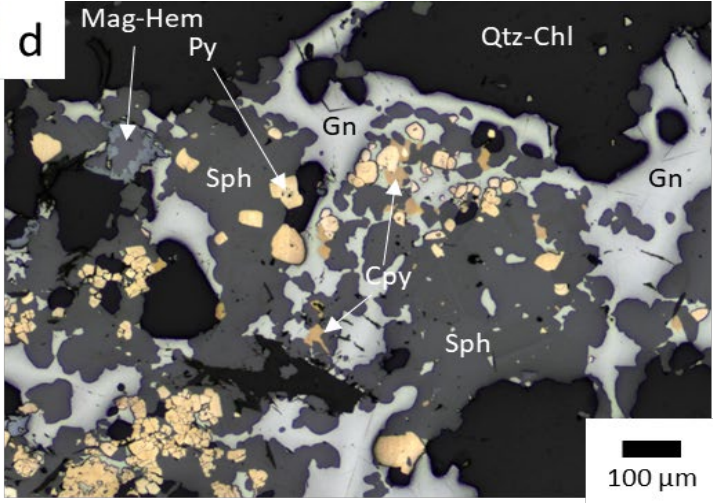
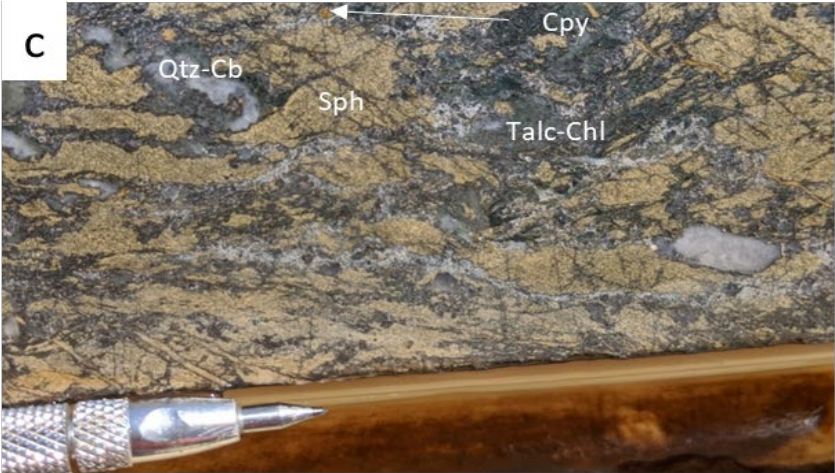
Bornite and chalcocite rims on pyrite grains in host rock



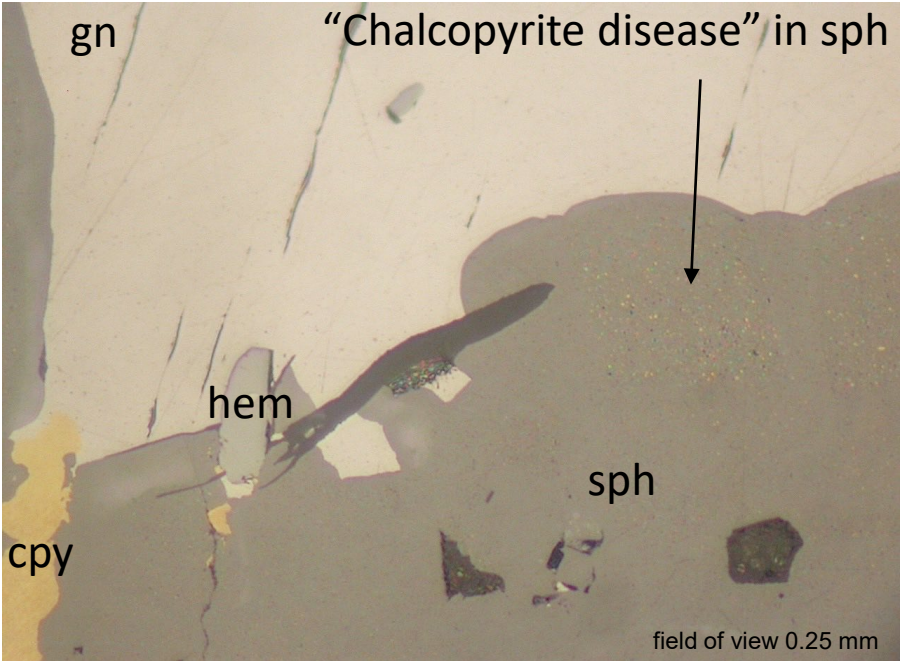
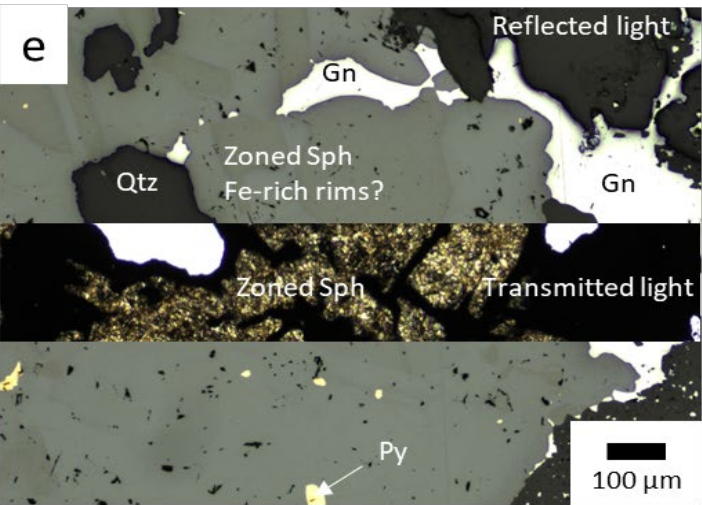
Cu-sulfide stringer veins below main Pb-Zn zone.



# Explorer 108 mineralisation

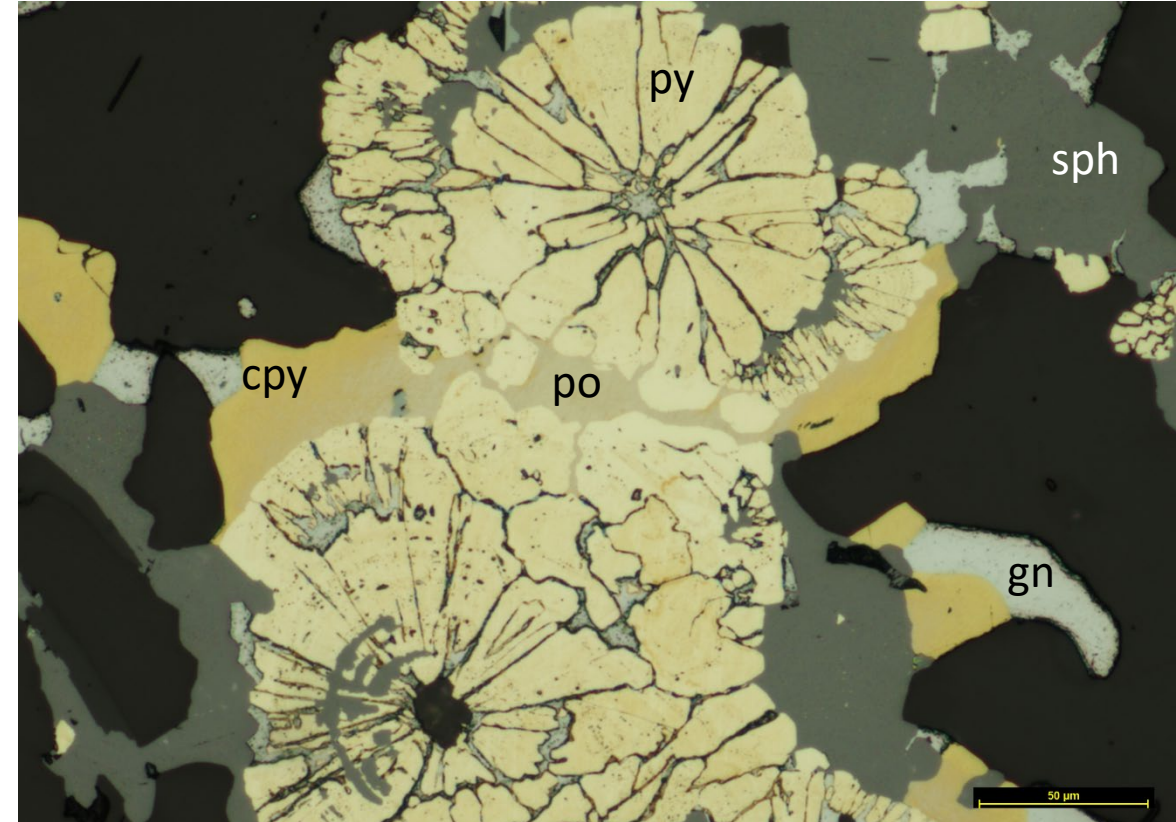
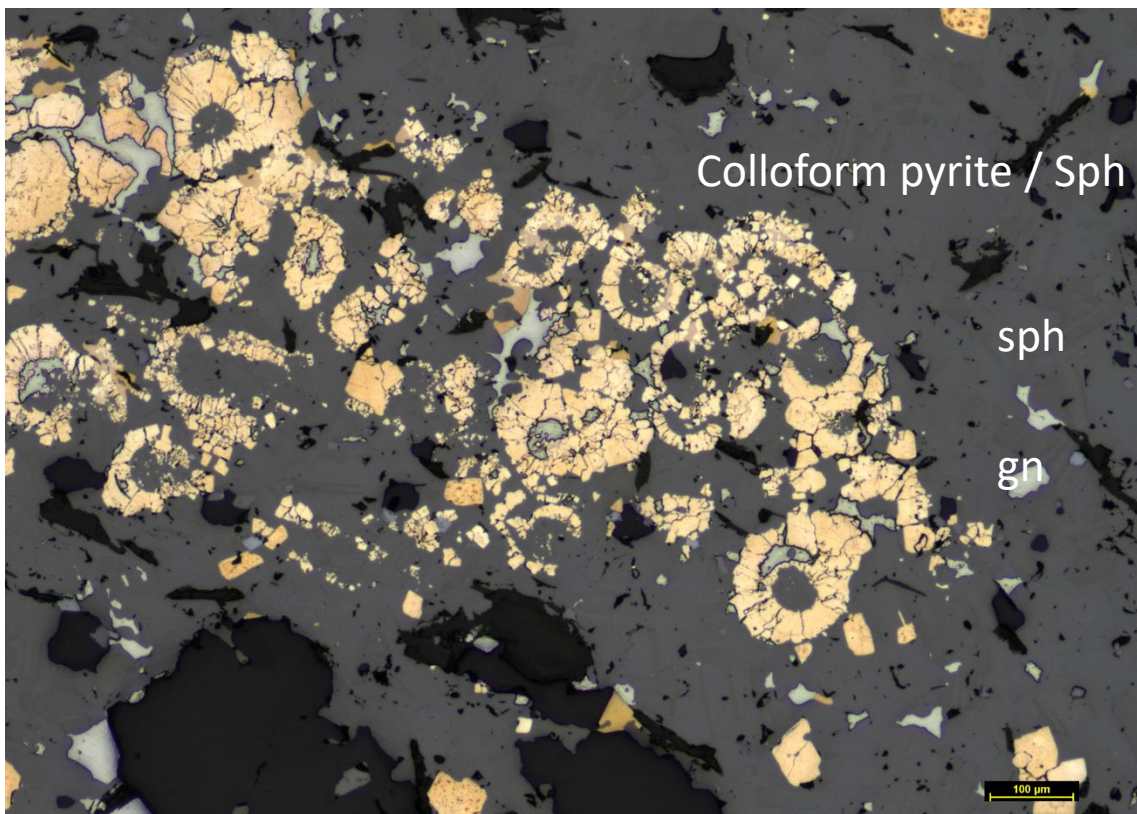


Exhalite beds? Concordant with Gn-Sph-rich layers



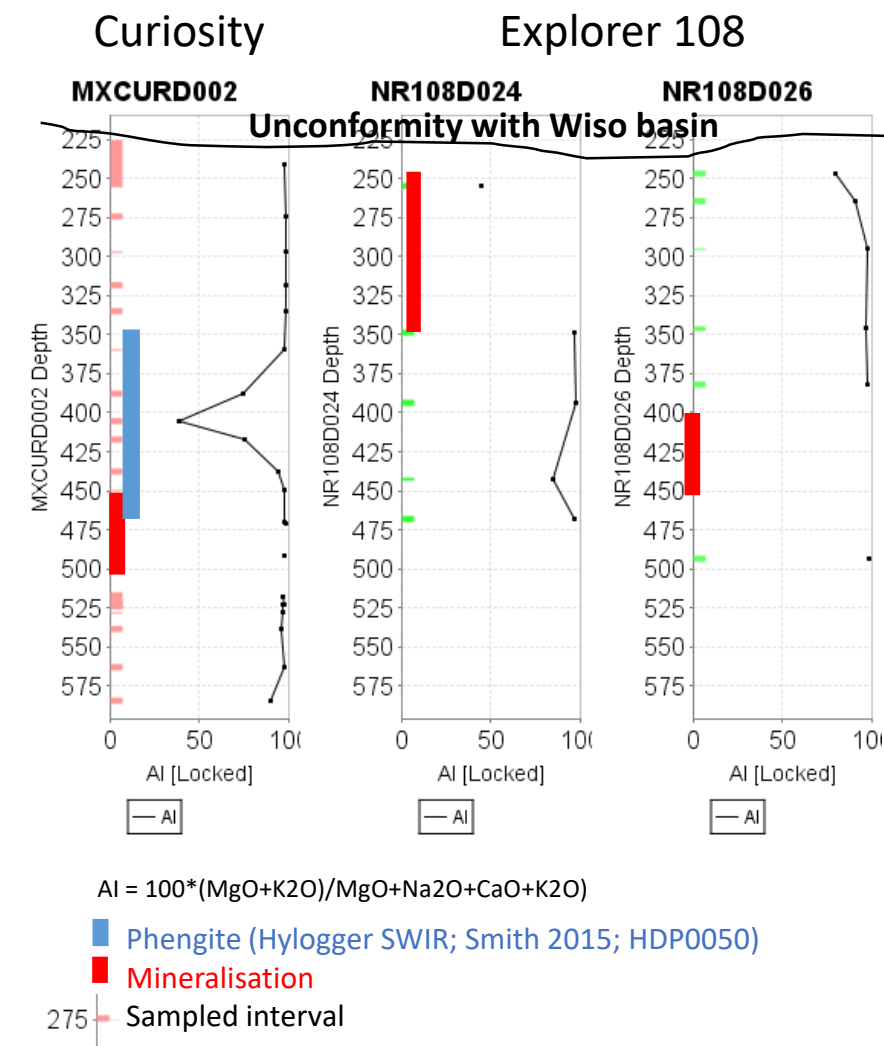
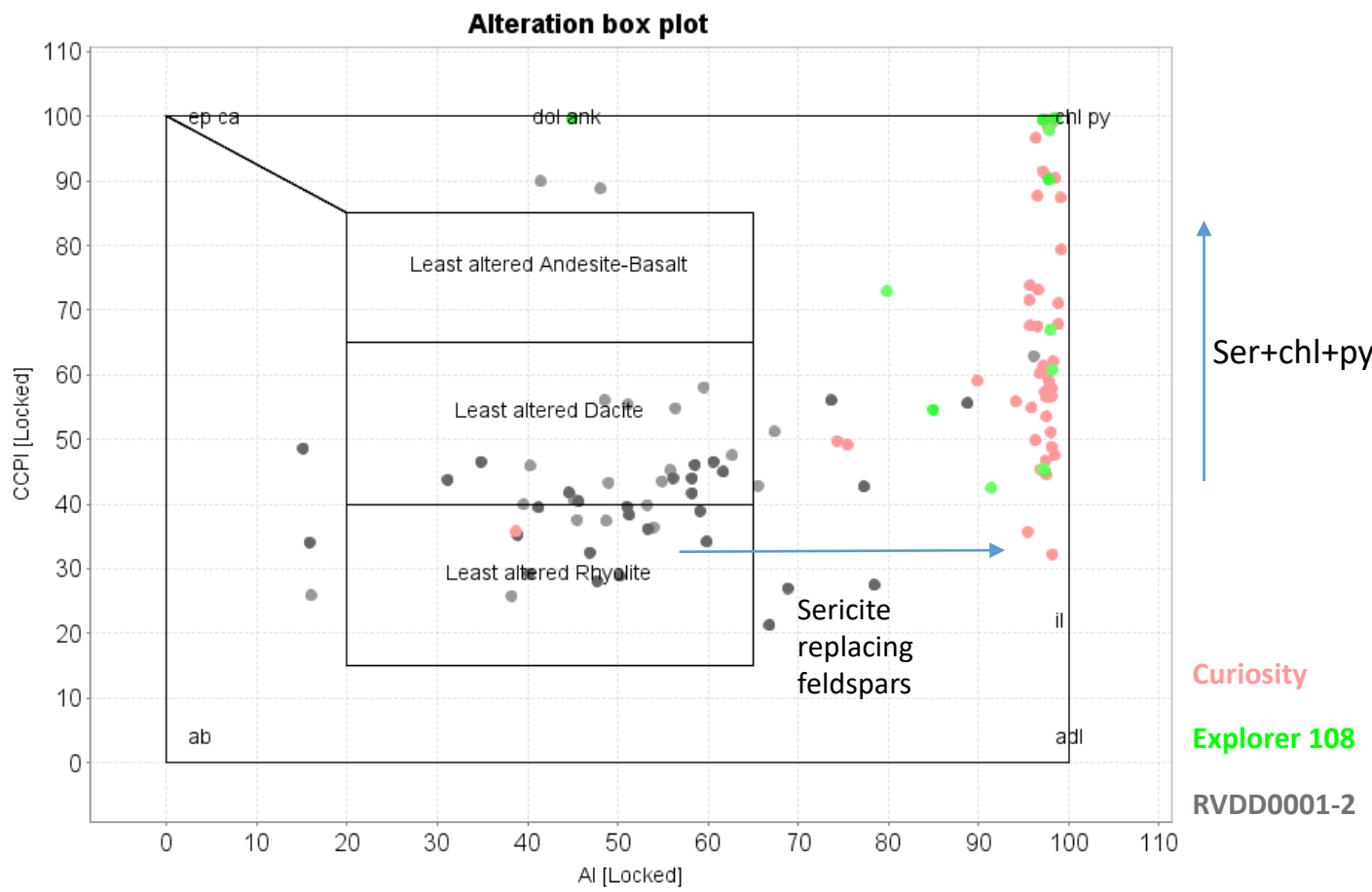
- Pb-Zn-Cu sulphides are associated with magnetite, it is common for the latter to show replacement by hematite, suggesting that the base metal sulphides were deposited with hematite and inferring oxidising conditions (consistent with the occurrence of Fe-poor sphalerite)
- Mineralisation associated with dolomite, talc and chlorite. All Mg bearing minerals





Aggregates of pyrite, locally sub-radiating overgrown by sphalerite (mid grey) and having apparent local replacement by pyrrhotite (pale creamy brown) and chalcopyrite (yellow).

# Explorer 108 / Curiosity alteration footprint



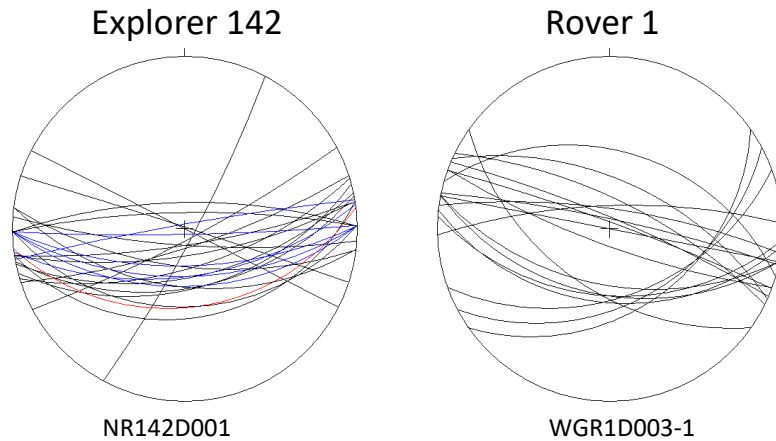
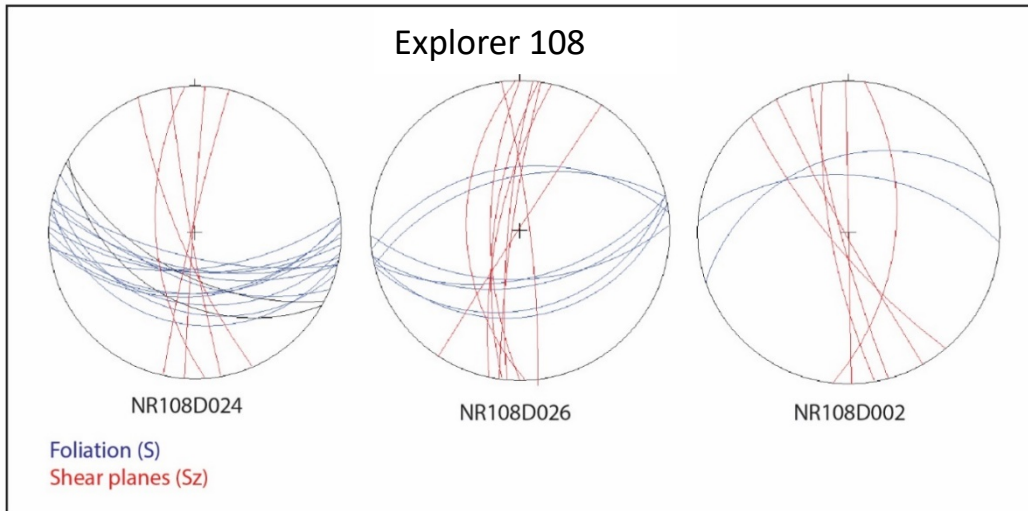
Chlorite Carbonate Pyrite Index (CCPI) vs Ishikawa alteration index (AI).

$$CCPI = 100 * (MgO + FeO) / (MgO + Na_2O + FeO + K_2O),$$

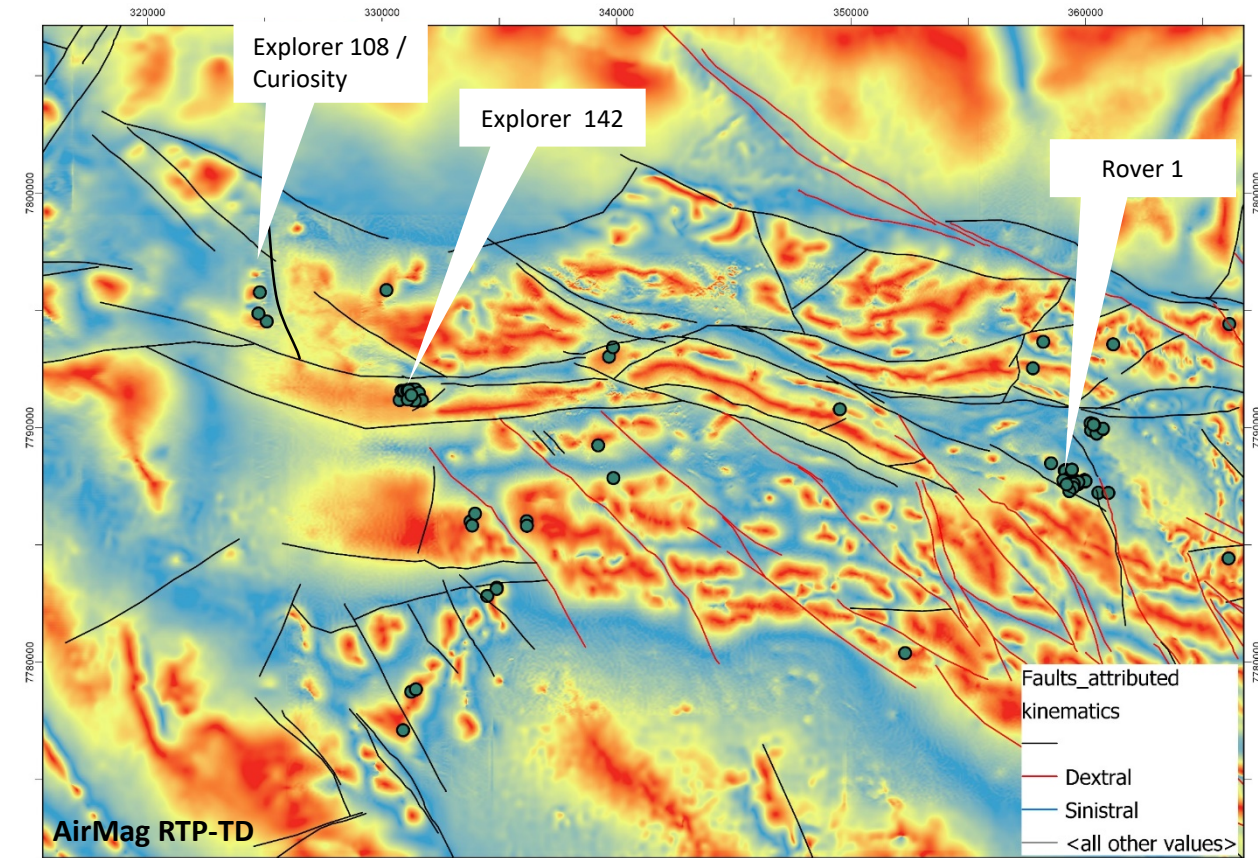
- Large alteration footprint! (~200m)



# Explorer 108 structures



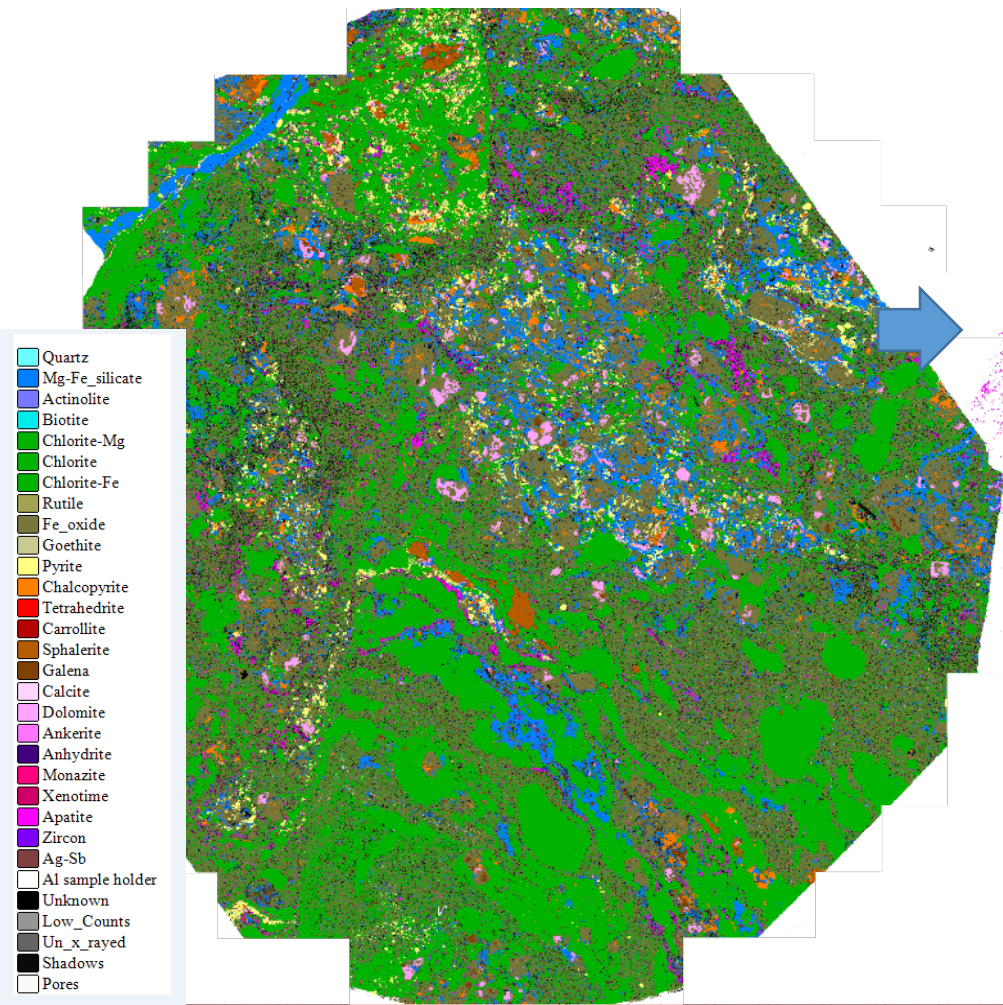
- Foliation is similar to the axial planar foliation in Rover 1 and Explorer 142
- N-S trending shear zones overprinting foliation. This shearing seems to be syn- to post-mineralisation



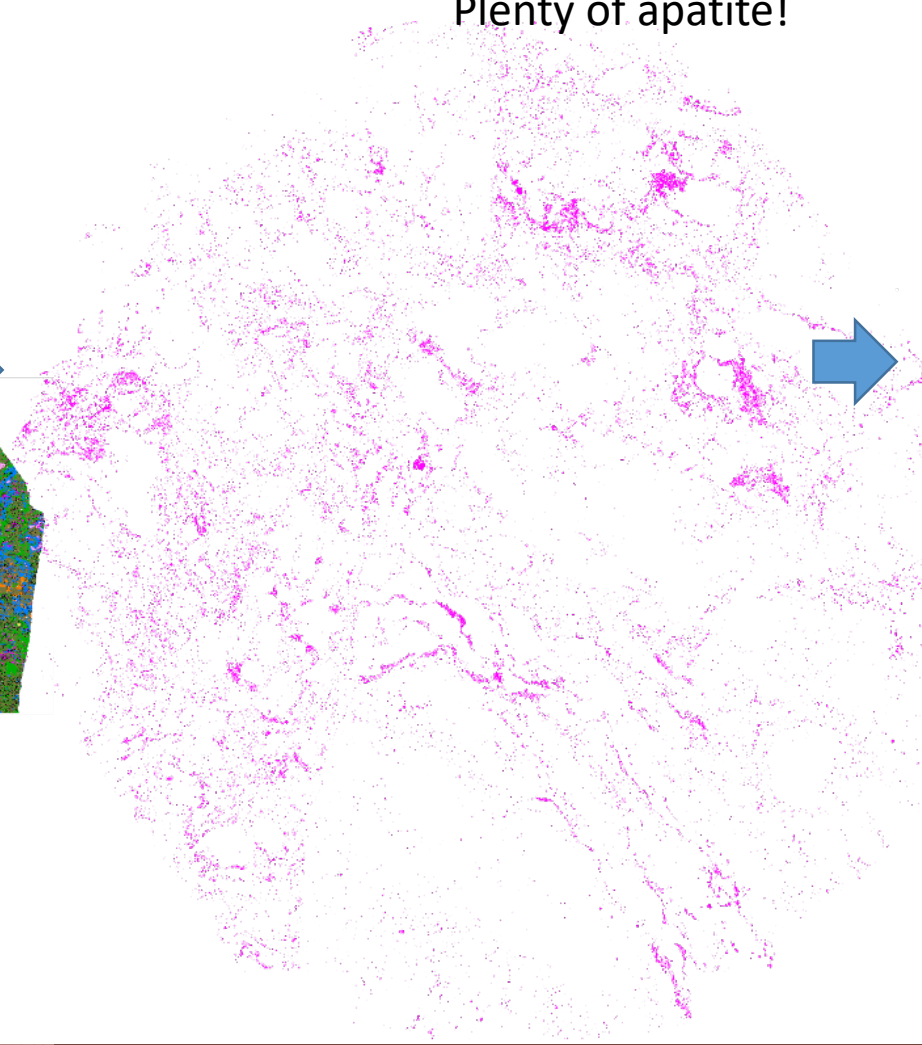


# Curiosity apatite U-Pb geochronology

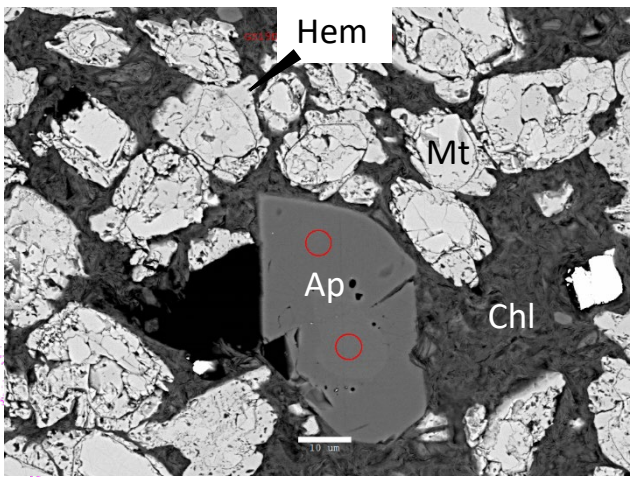
Mineral Liberation Analyser (MLA) image



Plenty of apatite!



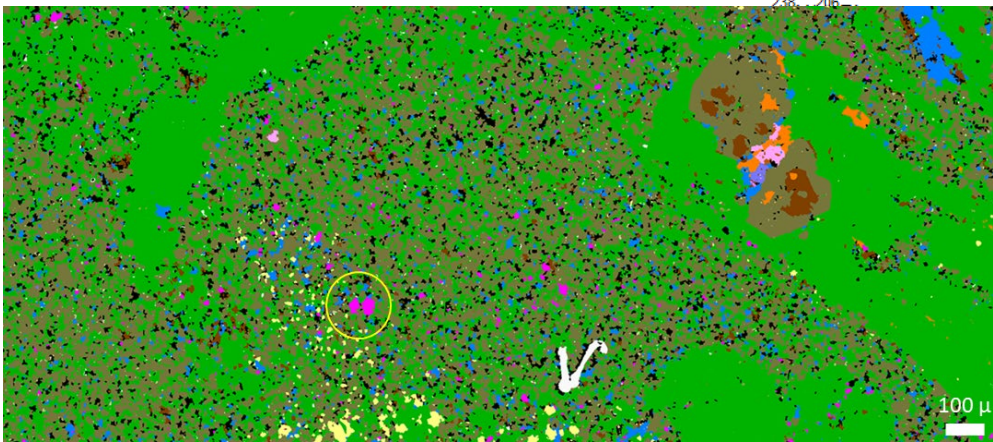
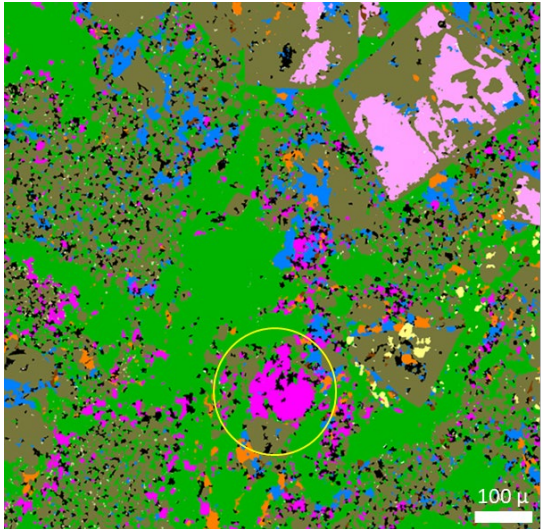
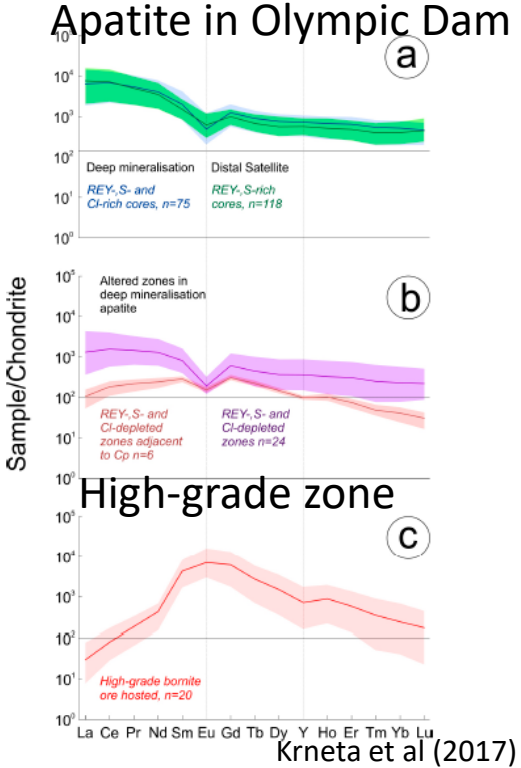
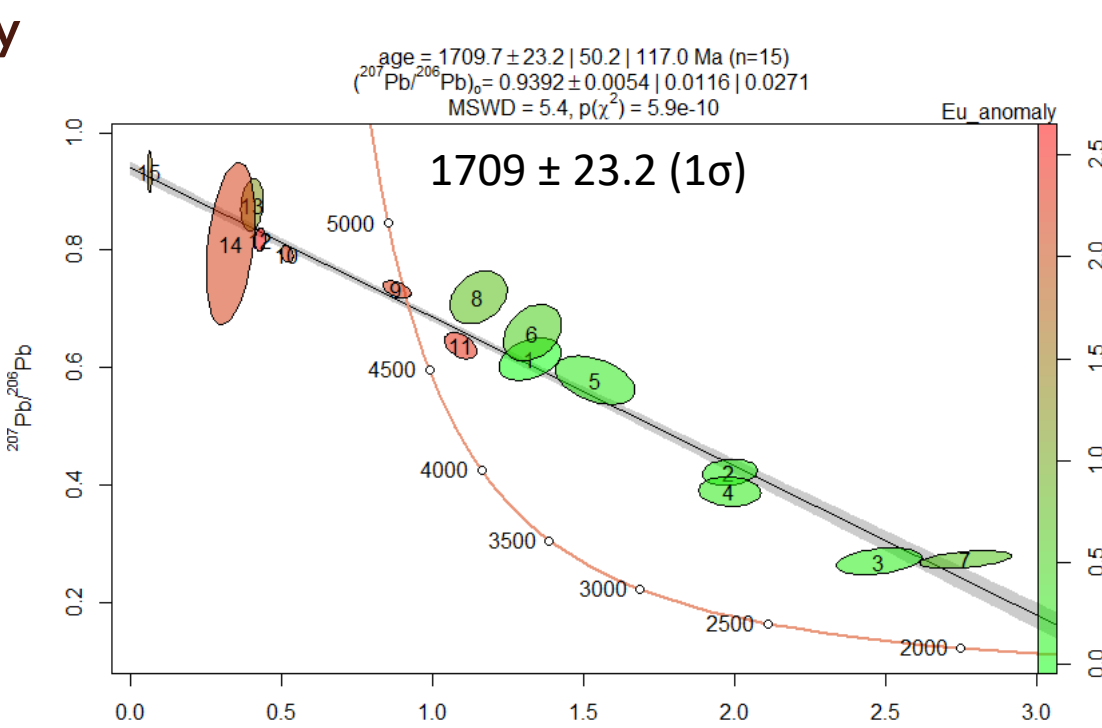
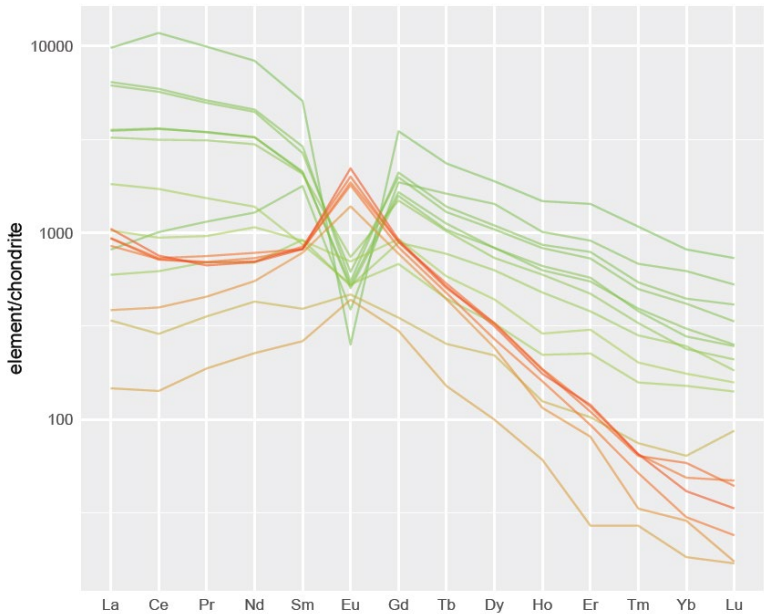
SEM image + EPMA



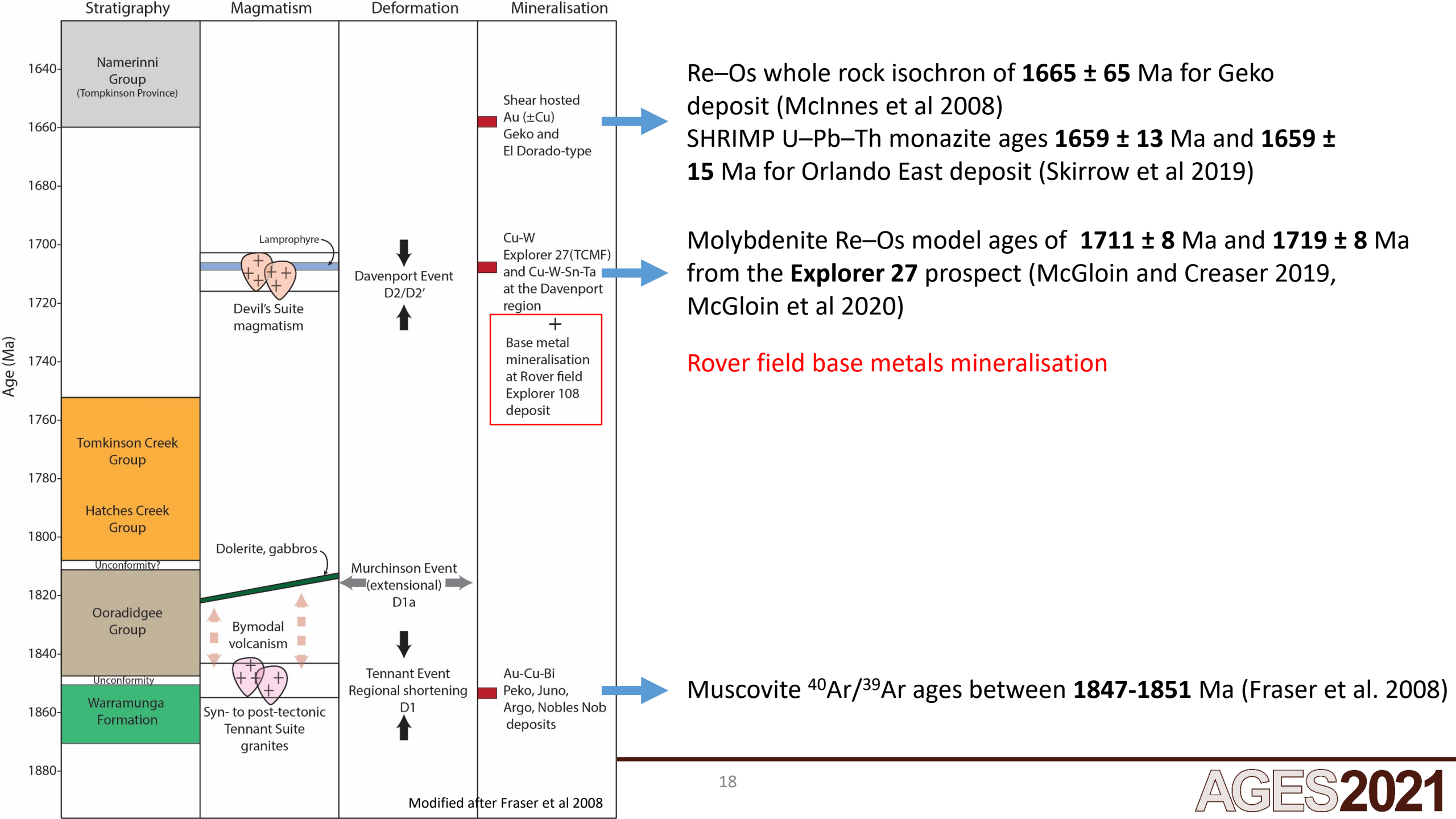
EPMA results:  
Fluorapatite  $\text{Ca}_5(\text{PO}_4)_3\text{F}$



# Curiosity apatite U–Pb geochronology



- + **Eu anomaly**= Large anhedral apatite grains / agglomerate hosted in Chl-Cpy-rich zones with large magnetite porphyroblasts
- **Eu anomaly**= Small 30–50 micron euhedral apatite grains hosted in fine-grained magnetite clusters / boudins





# Conclusions

## About the host of mineralisation in the Rover field

- Rover field host stratigraphy seems to be similar (in age and lithology) to the **Yungkulungu Fm** (Ooradidgee Gr)
- The volcanoclastic rocks range in composition **from andesitic to rhyolitic**
- Rover 1, Explorer 142 and Curiosity are hosted in turbidites. Explorer 108 is hosted in more immature proximal volcanoclastic facies.

## About the base metal mineralisation

- Alteration footprints ( $AI > 75$ ) are broad and reach the unconformity with the Wiso Basin
- Host rocks, sulfide assemblages and textures similar to that of VHMS systems (eg Cu-rich stringer zones, exhalites, colloform pyrite, zone refining; Galley et al 2007, 2015; Large 1992; Hollis et al 2015). **However** the age of mineralisation (~1700 Ma) in comparison with the age of the host (~1850 Ma) suggest epigenetic mineralisation

## Age of host and mineralising event

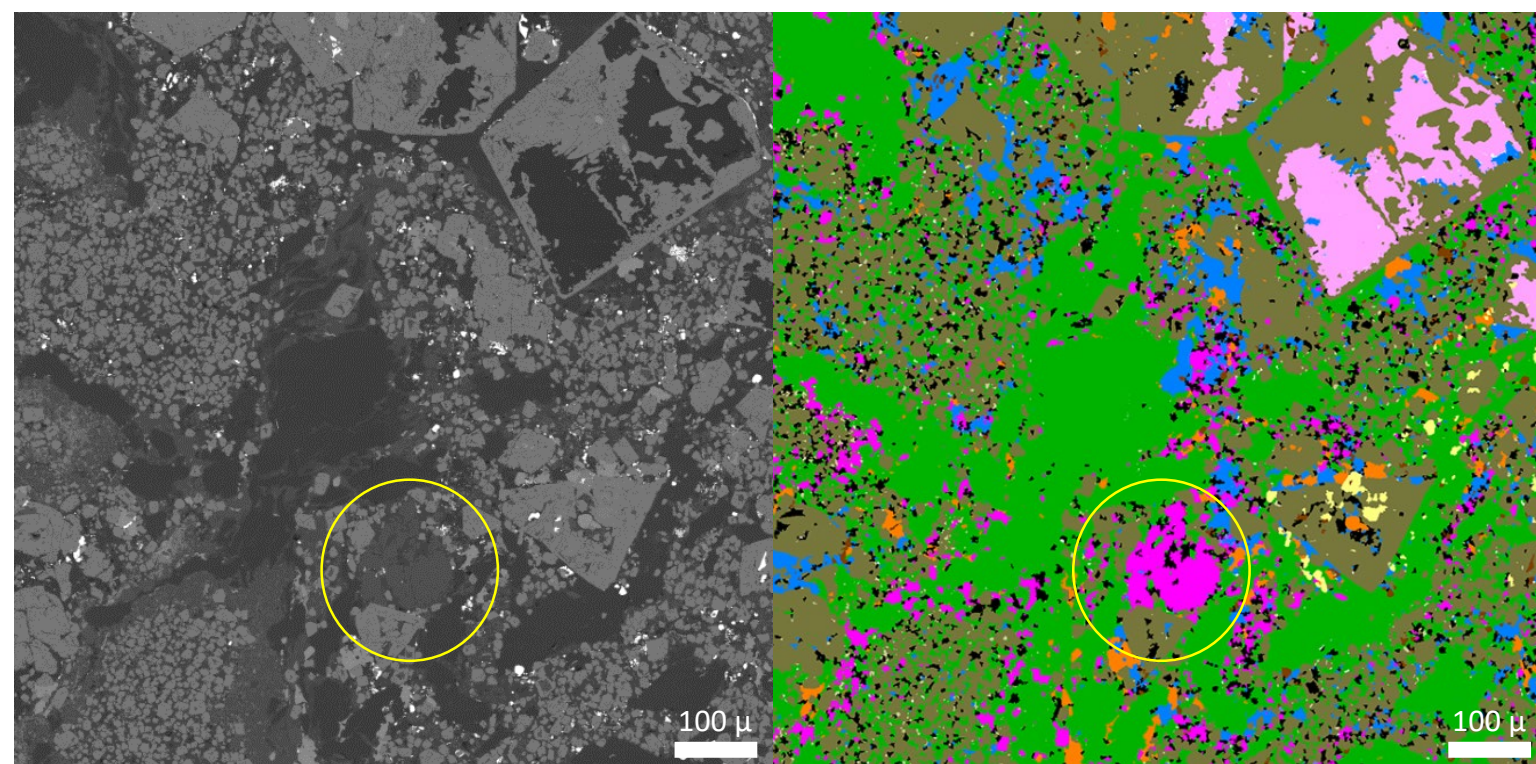
- Host MDA and crystallisation ages suggest Yungkulungu Fm
- Apatite age indicates mineralisation contemporary with the Davenport Event, Devil's Suite magmatism and other Cu-W mineralisation events

THANK YOU

QUESTIONS?



# Curiosity apatite U-Pb geochronology

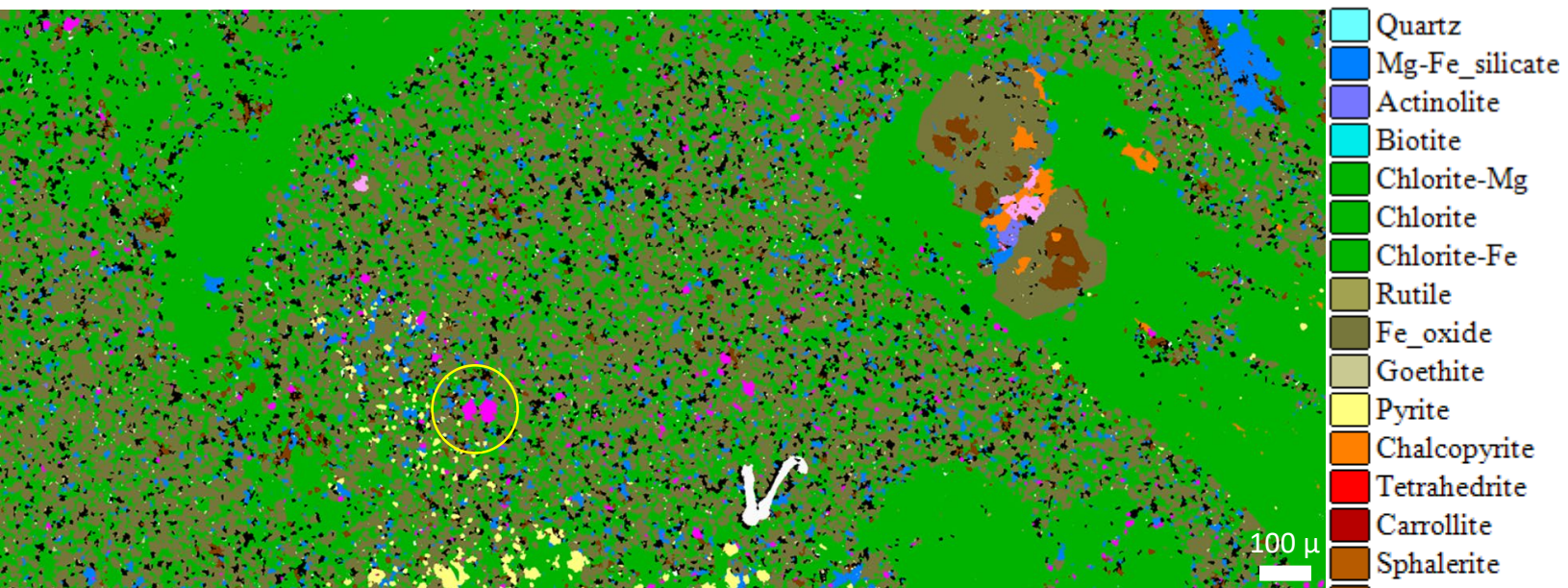


- Quartz
- Mg-Fe\_silicate
- Actinolite
- Biotite
- Chlorite-Mg
- Chlorite
- Chlorite-Fe
- Rutile
- Fe\_oxide
- Goethite
- Pyrite
- Chalcopyrite
- Tetrahedrite
- Carrollite
- Sphalerite
- Galena
- Calcite
- Dolomite
- Ankerite
- Anhydrite
- Monazite
- Xenotime
- Apatite
- Zircon
- Ag-Sb

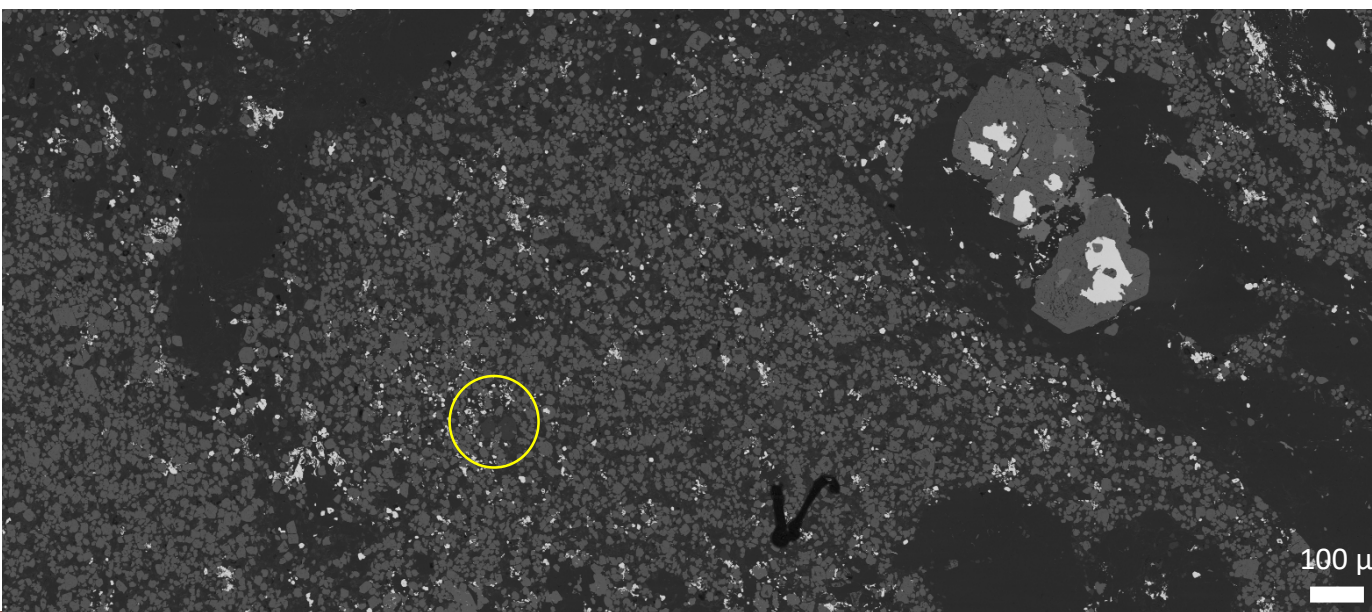
+ **Eu anomaly**= Large anhedral apatite grain / agglomerate hosted in Chl-Tlc-Cpy-rich zones with large magnetite porphyroblasts



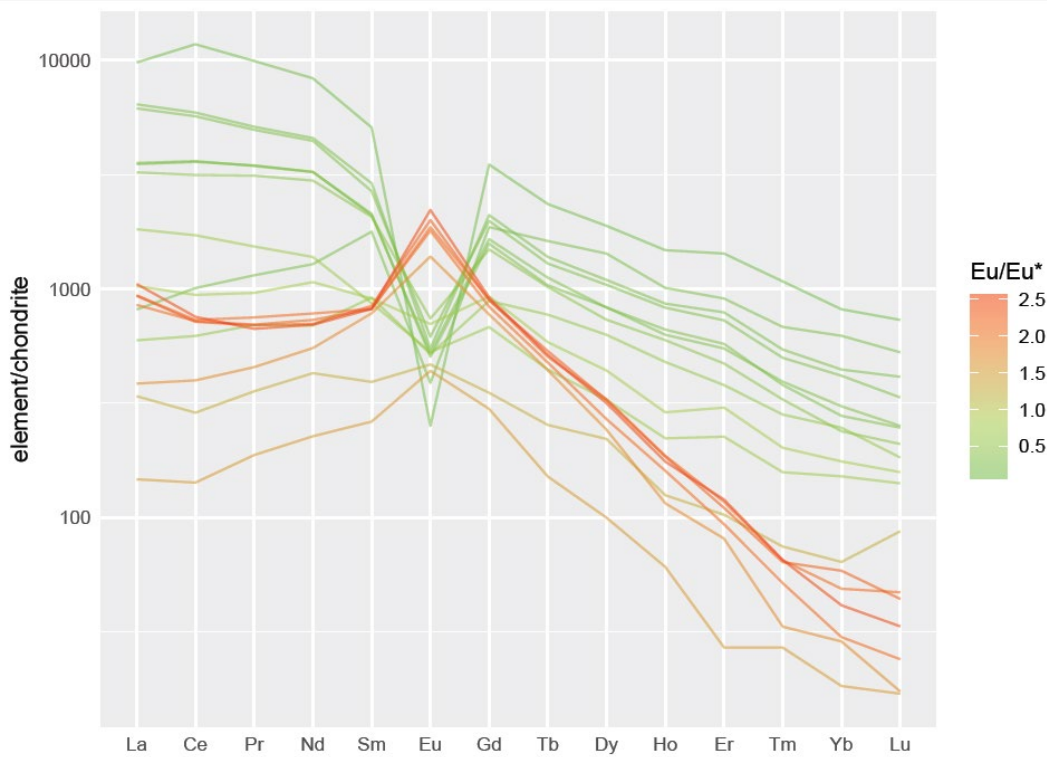
# Curiosity apatite U-Pb geochronology



- **Eu anomaly**= Small 30–50 micron euhedral apatite grains hosted in fine-grained magnetite clusters / boudins

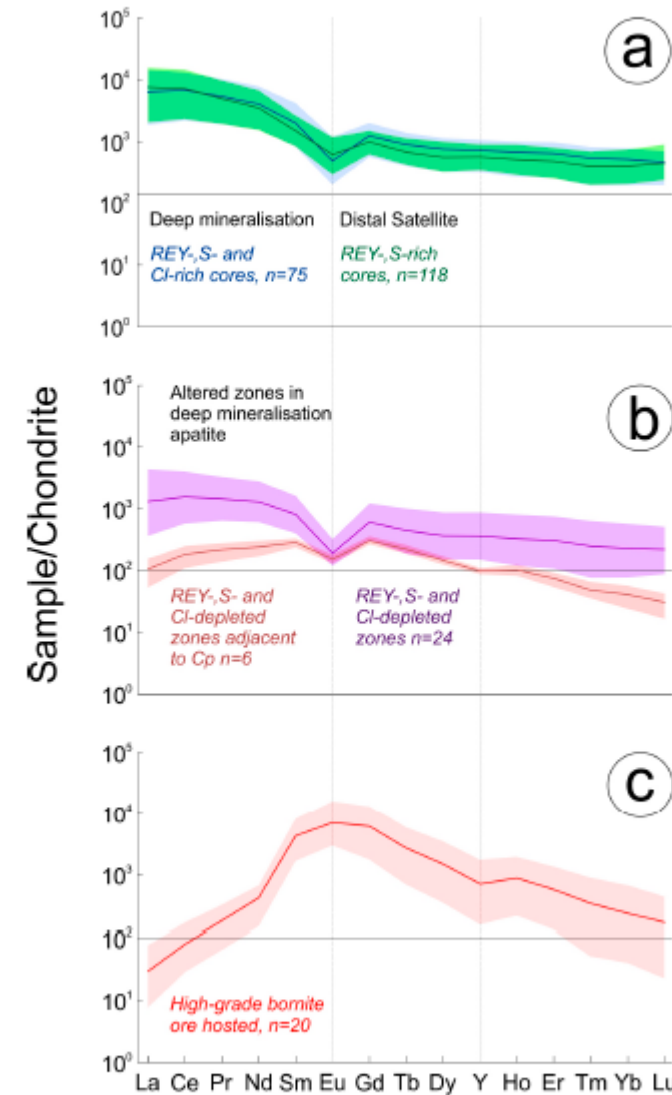






## Hydrothermal apatite

## Apatite in Olympic Dam



Krneta et al (2017)

**Figure 5.** Chondrite-normalized [44] LA-ICP-MS plots of hydrothermal apatite. Analyses were performed on multiple grains hosted within samples listed in the captions to the respective tables. (a) Hydrothermal apatite hosted in the Deep and Distal Satellite Mineralization. (b) Altered zones in Deep Mineralization hosted apatite related. (c) Apatite hosted within the high-grade bornite ore.

So, what type of deposit is Explorer 108 and Curiosity?

Mt Isa SEDEX (Savage 2020)

“In the least deformed Australian deposits (Halleyer and Mount Chalmers) many textural features have been observed which are similar to those recorded for the undeformed kuroko deposits by Barton (1978) and Eldridge et al. (1983). Such features include growth-zoned sphalerite, **chalcopyrite-diseased sphalerite, colloform pyrite, and colloform sphalerite-pyrite intergrowths.**” (Large, 1992).

Zone refining?

Sph overprinted by Cpy via coupled dissolution-reprecipitation

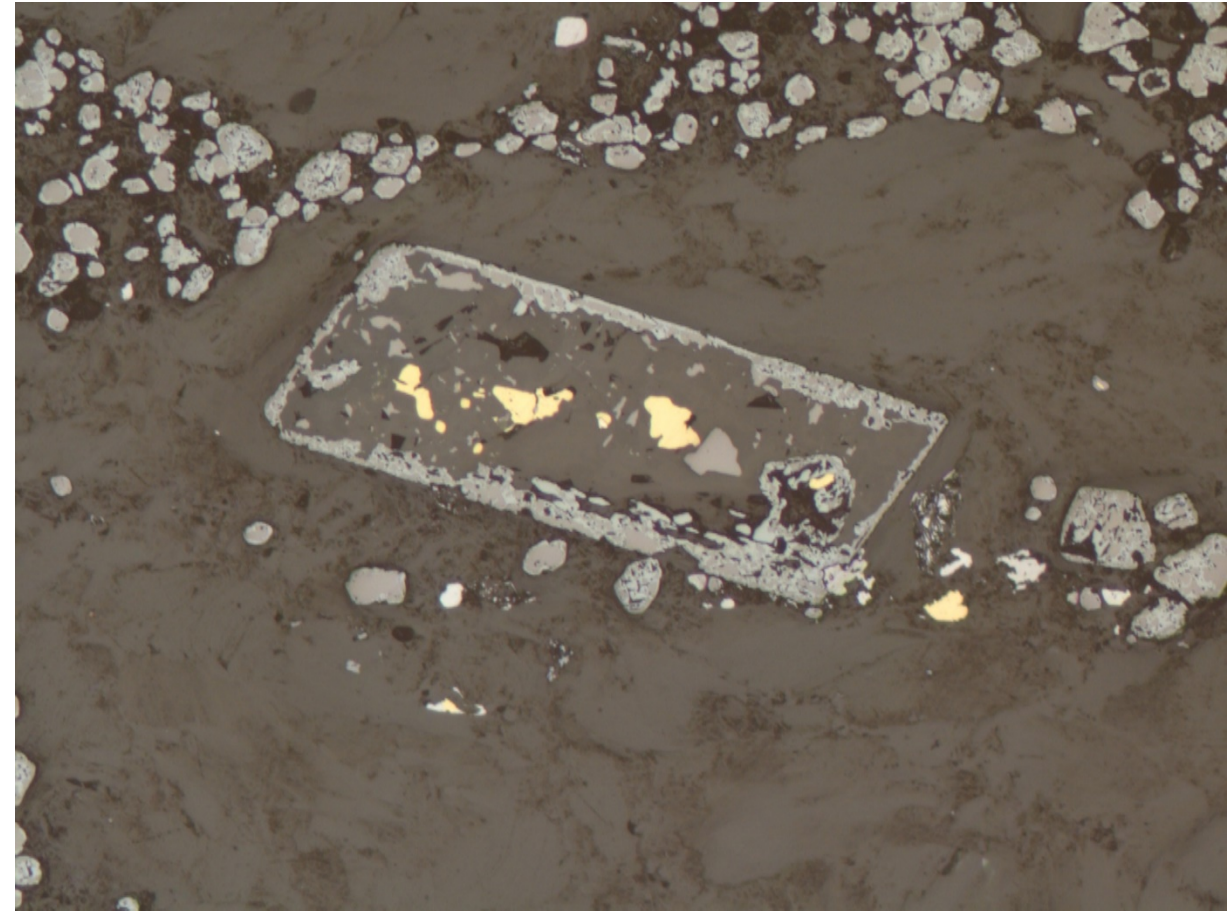
Hydrothermal influx of Mg and base metals sulphides under oxidising conditions (mt + Fe-poor sph + hem)



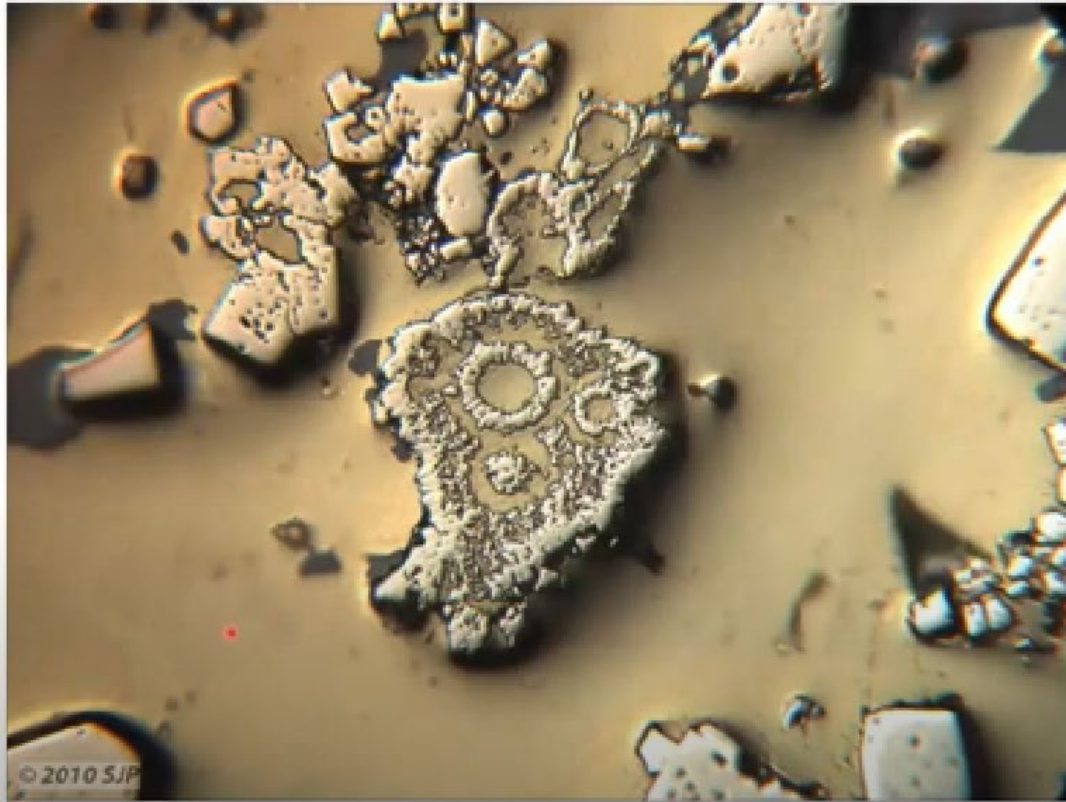
Could be



*Coarse (up to 2 mm) tabular high-relief barite crystals in finer grained interstitial quartz. Plane polarised light, field of view = 3.0 mm.*



Thank you!



Photomicrograph of colloform pyrite with chalcopyrite in ancient hydrothermal chimneys, Duck Pond, NL, Canada





